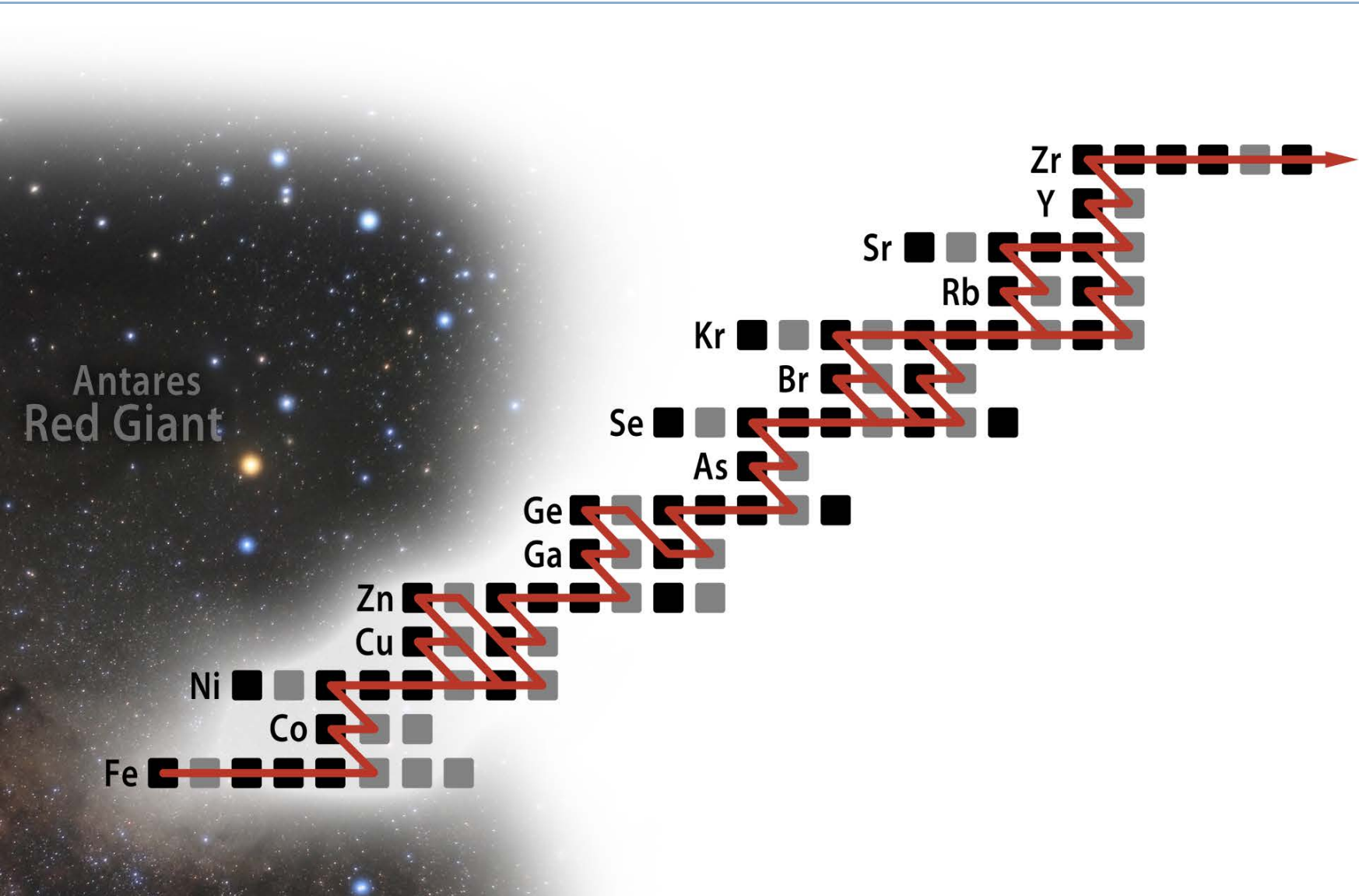


# Experiments for the astrophysical s-process

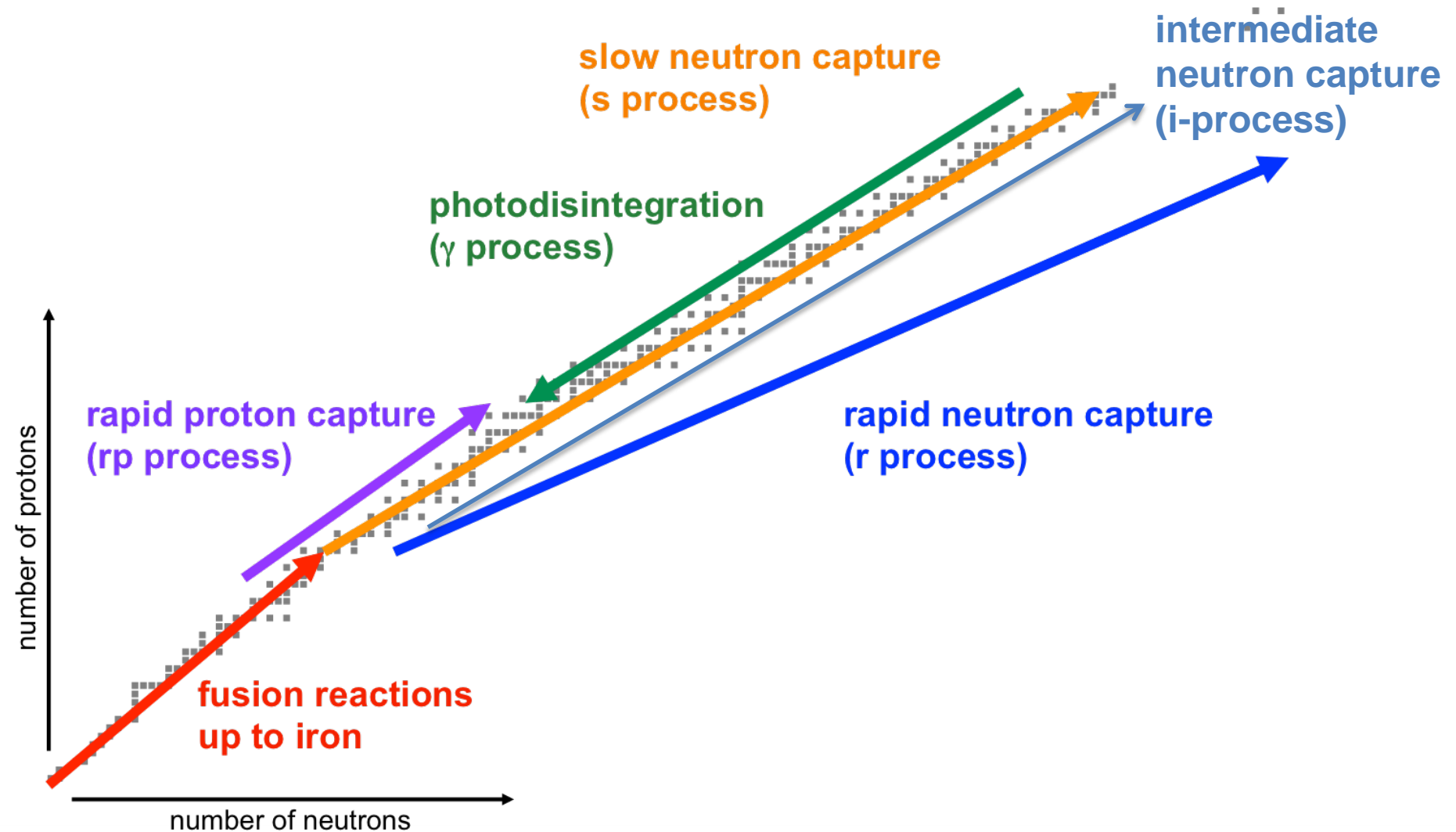
René Reifarth

*Workshop on Nuclear Astrophysics at the Dresden Felsenkeller*  
June 26, 2017, HZDR, Rossendorf, Germany

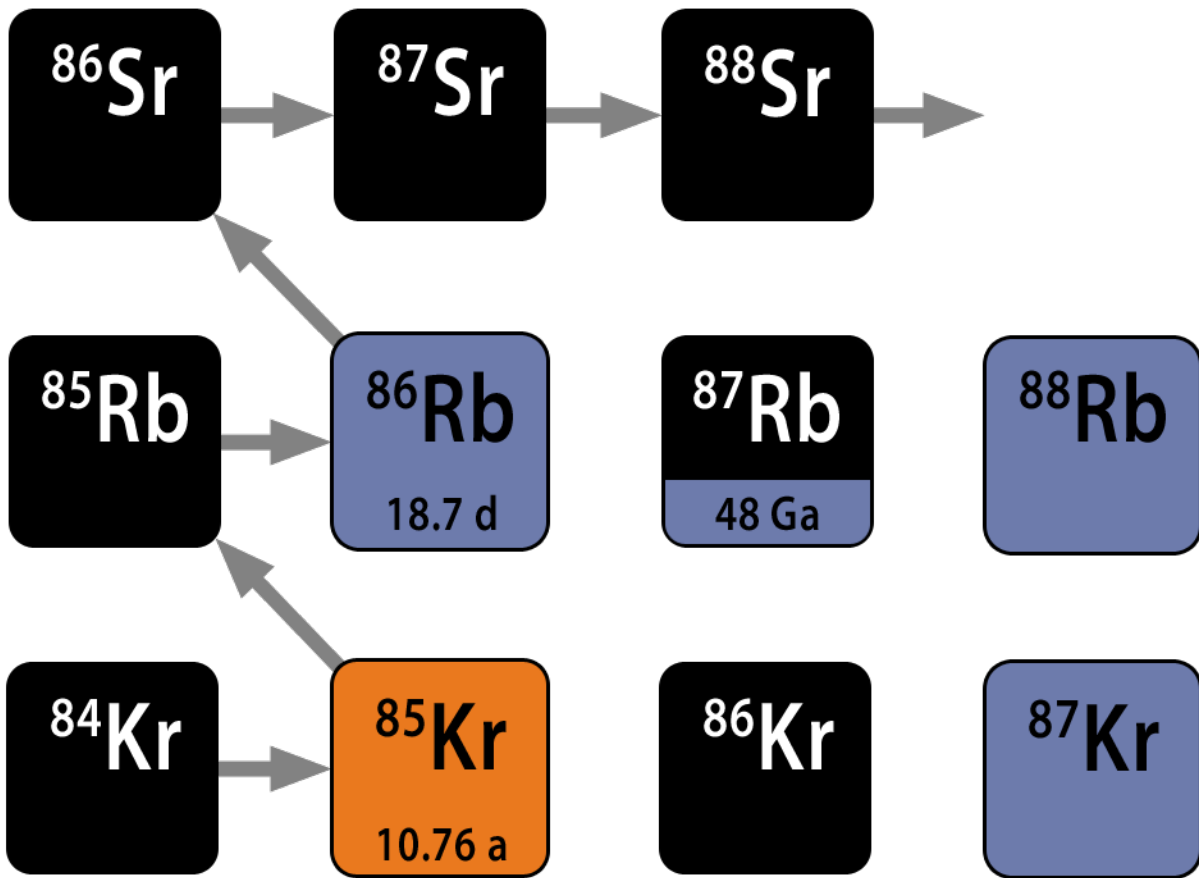
# Nucleosynthesis – tales from the past



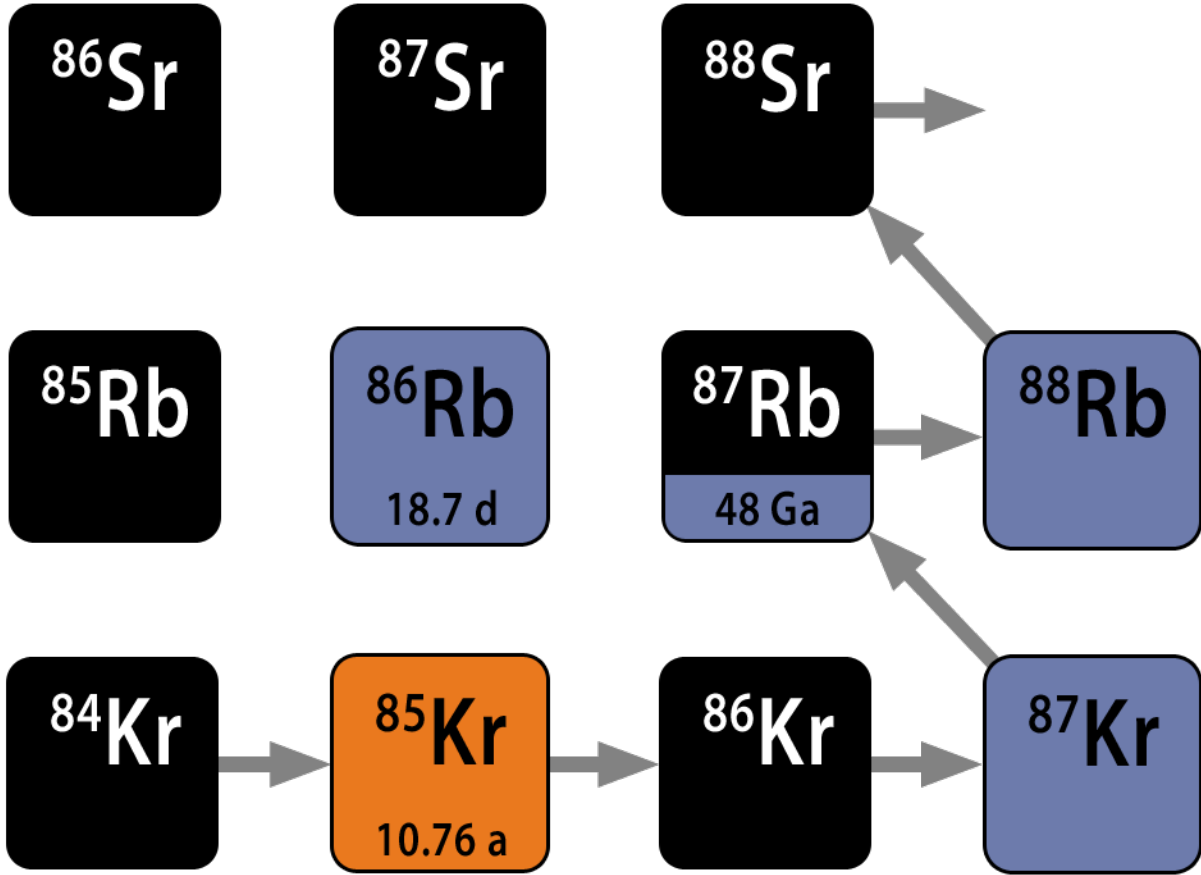
# The nucleosynthesis of the elements



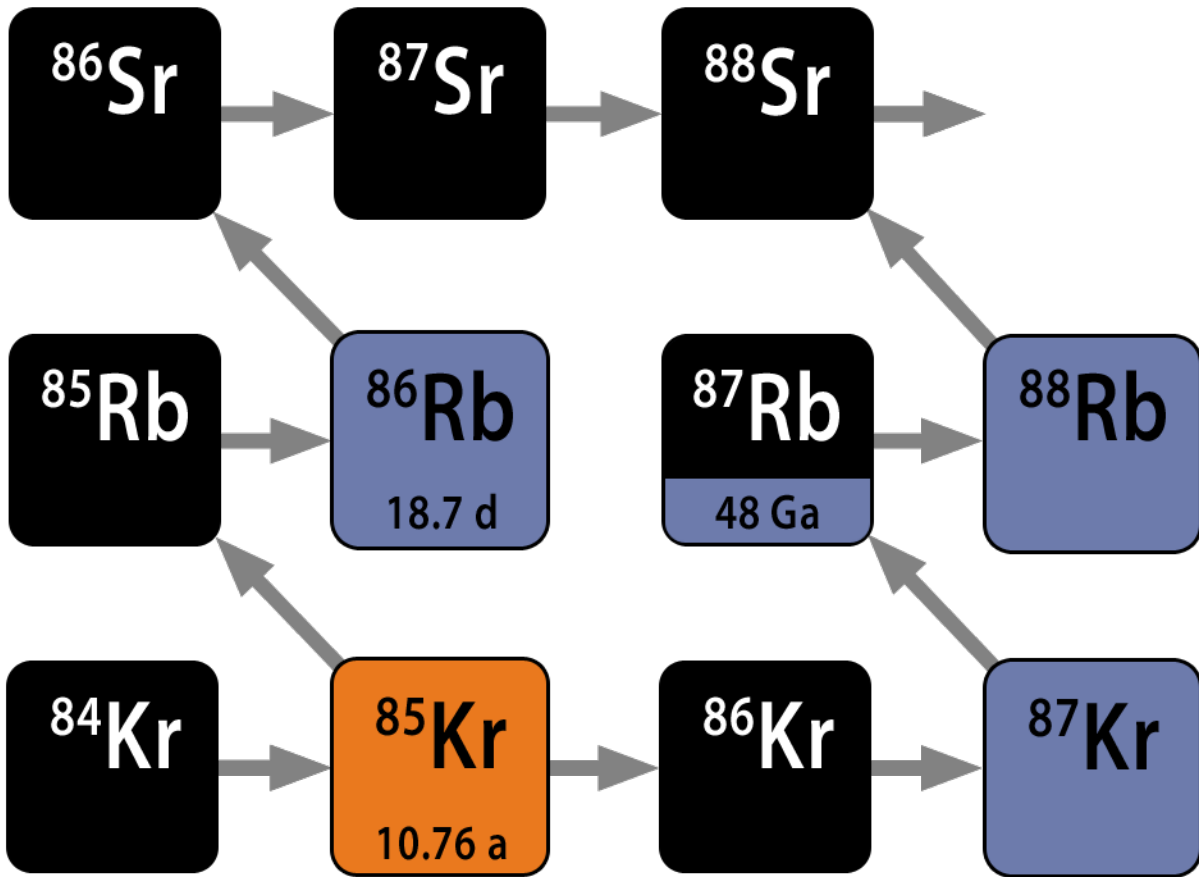
# Radioactive isotopes in the s-process



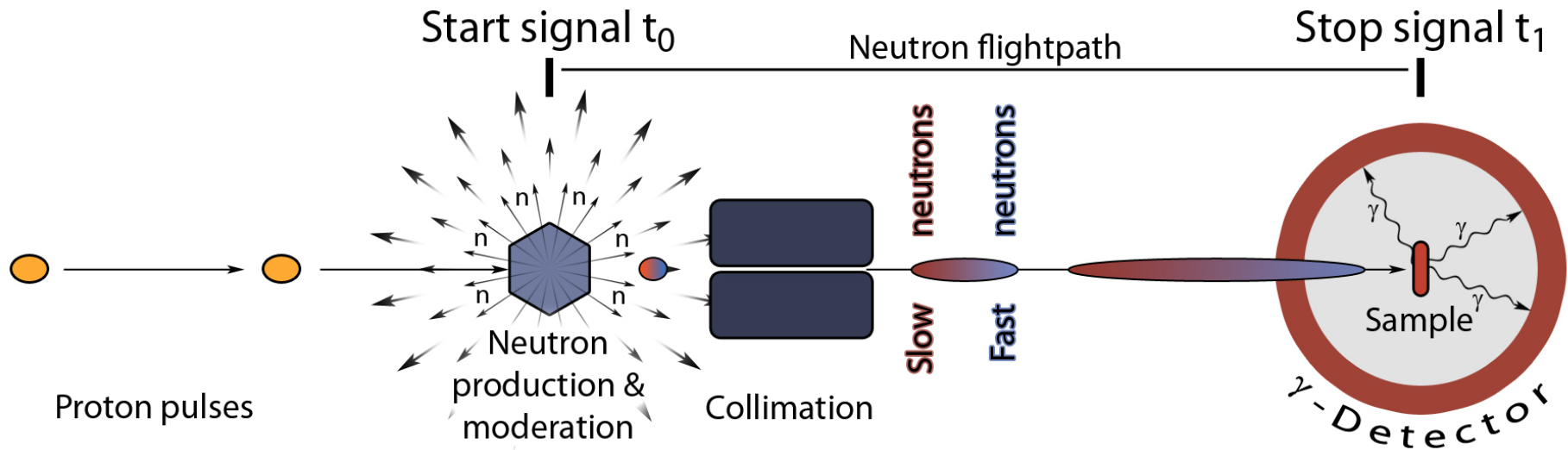
# Radioactive isotopes in the s-process



# Radioactive isotopes in the s-process

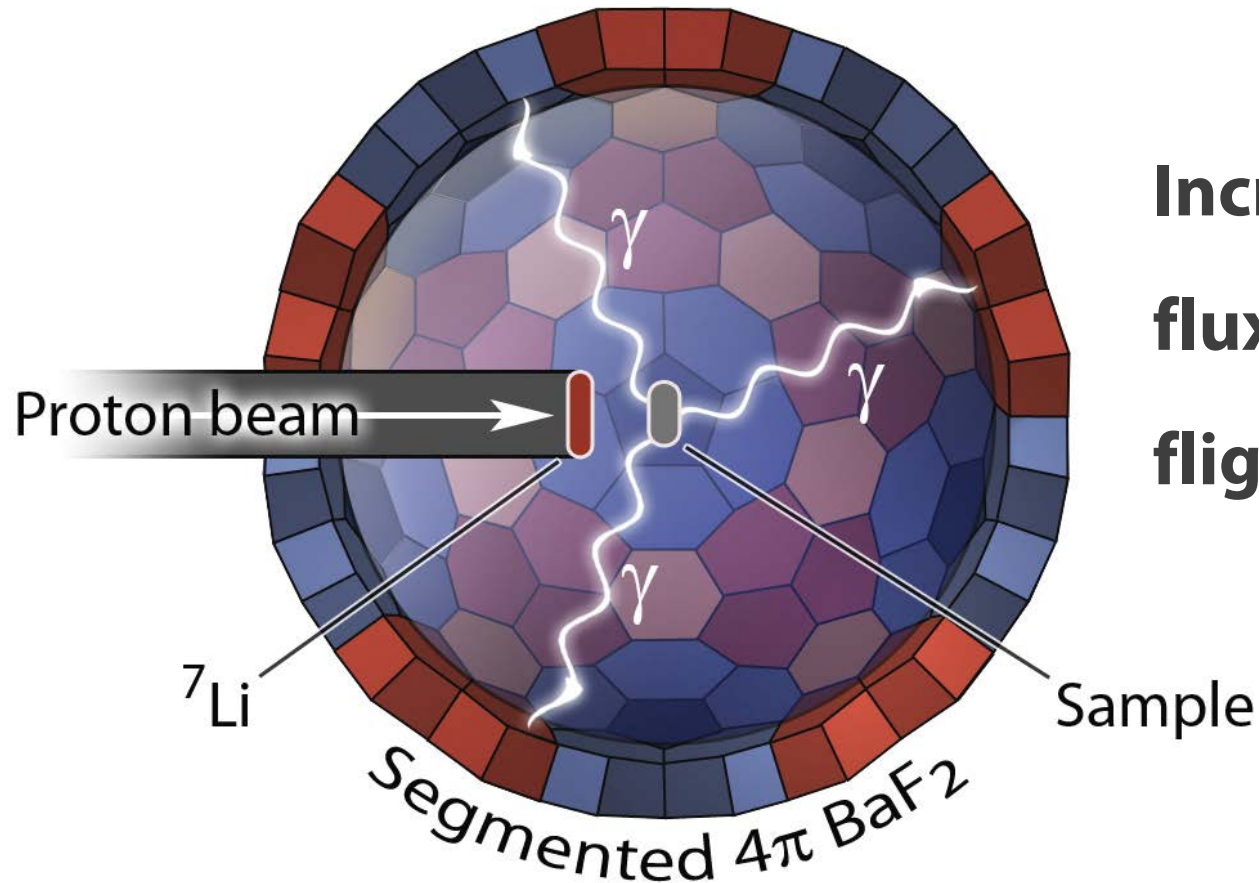


# Neutron Captures – time-of-flight technique



- the TOF-technique is the only generally applicable method to determine energy-dependent neutron capture cross sections
- beam pulsing & distance to the neutron production site significantly reduce the number of neutrons available on the sample

# NAUTILUS – Neutron capture with short flightpath

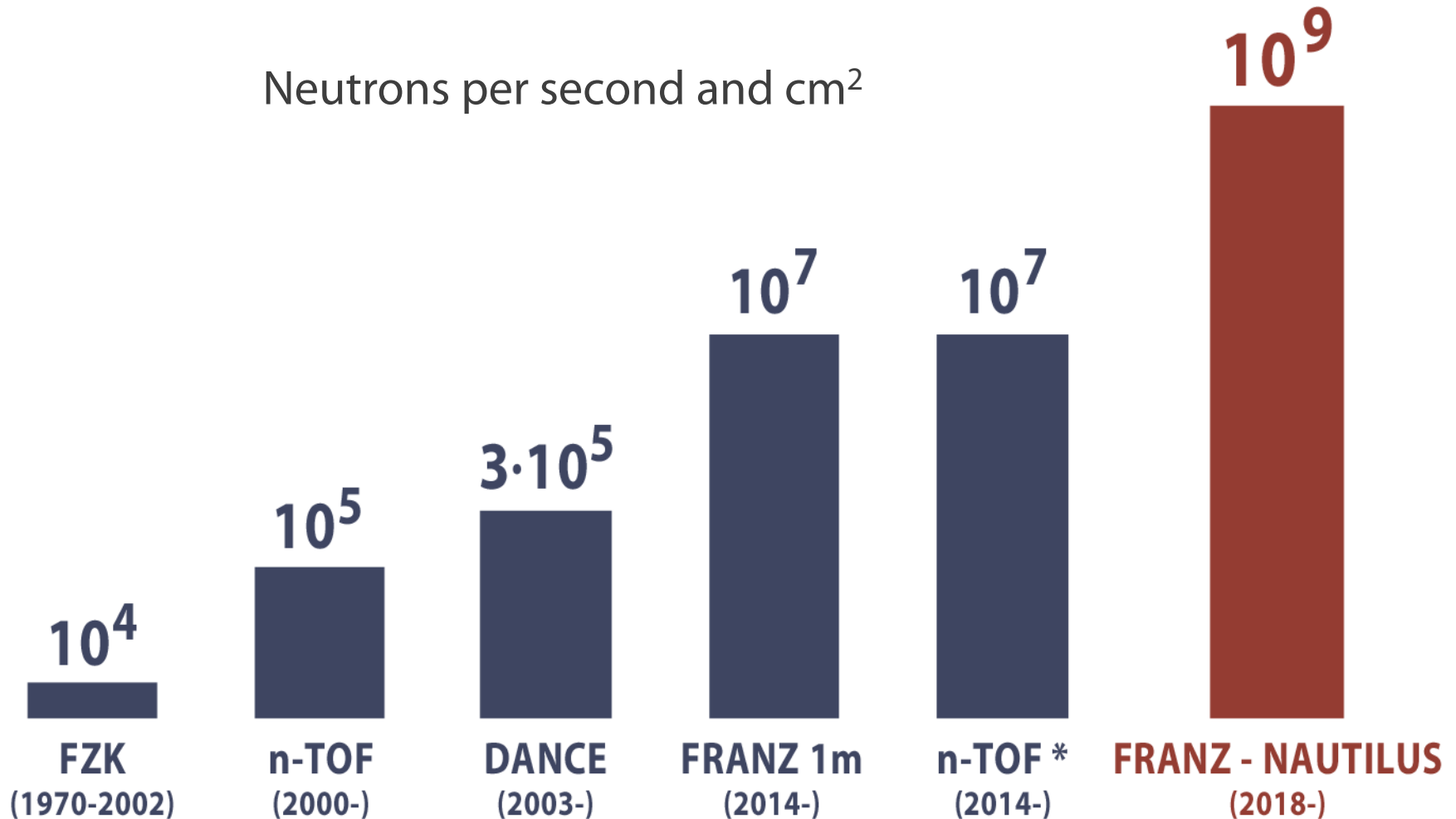


**Increase neutron  
flux with ultra-short  
flightpath**



# Neutron flux in astrophysical region

Neutrons per second and cm<sup>2</sup>



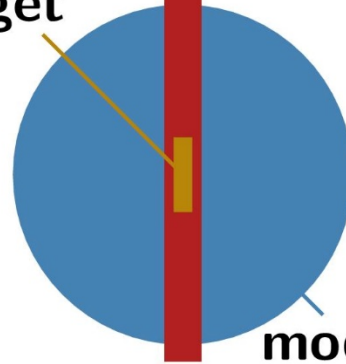
↓ Protons

Tungsten spallation target



↓ Protons

Tungsten spallation target



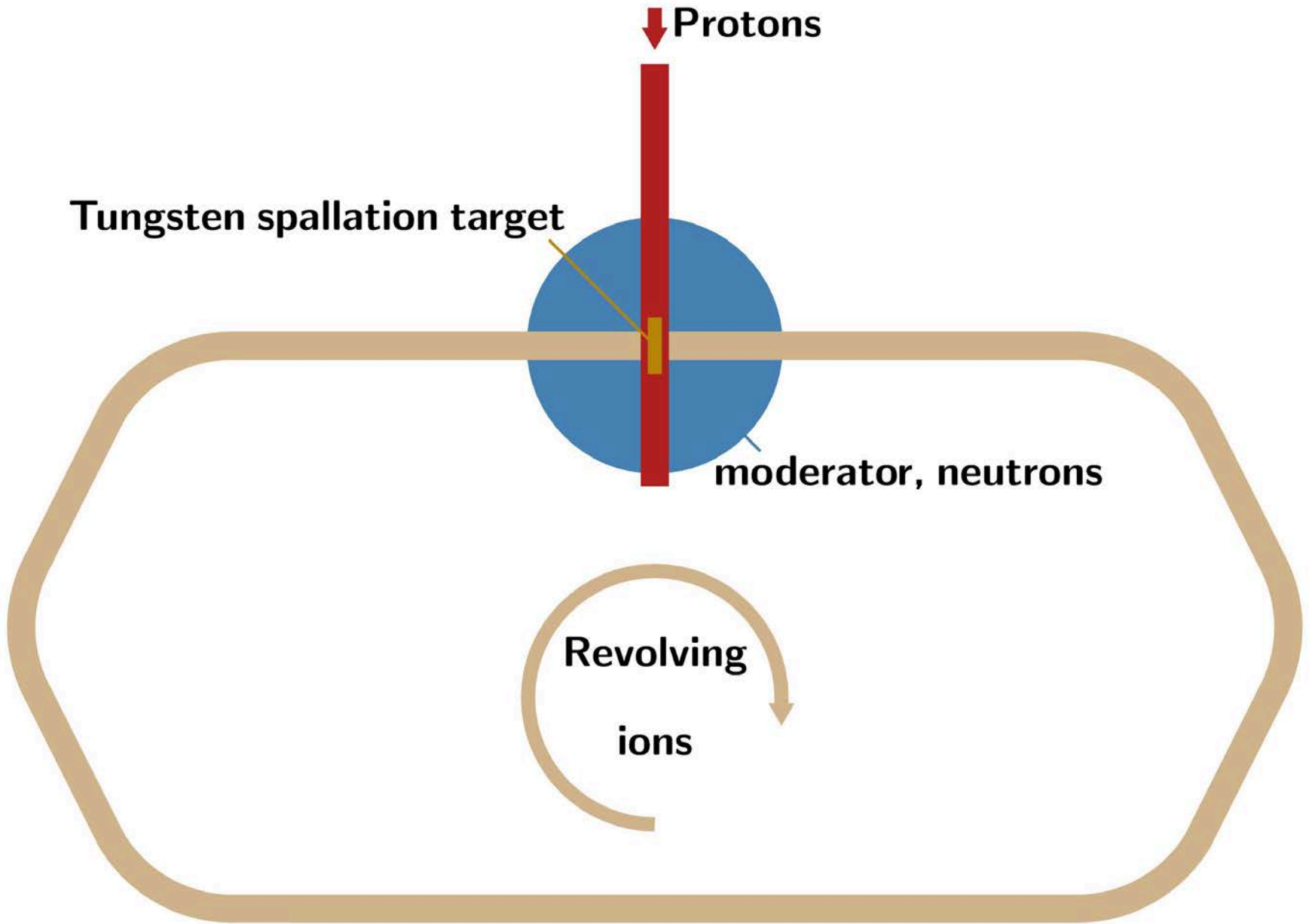
moderator, neutrons

↓ Protons

Tungsten spallation target

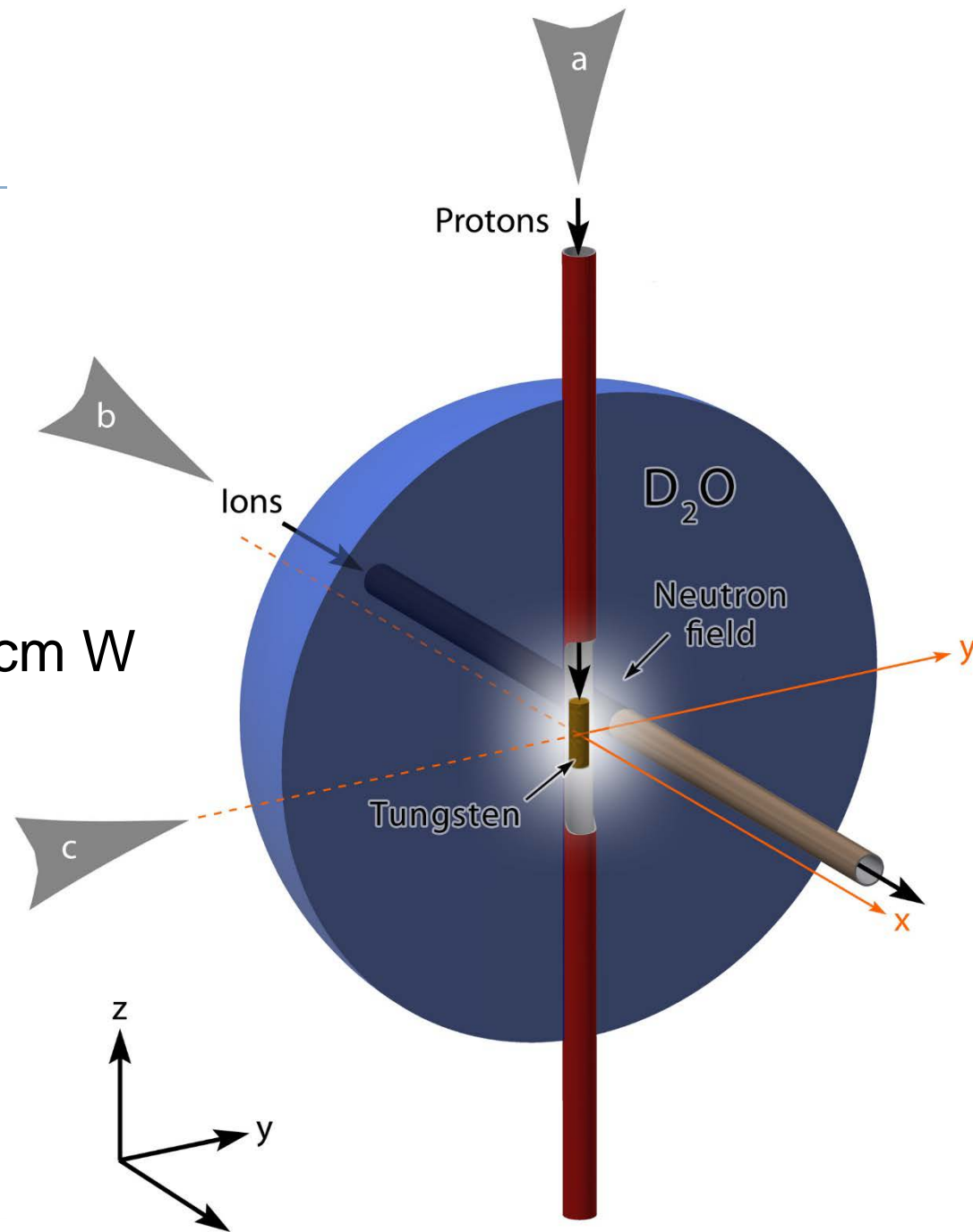
moderator, neutrons

Revolving  
ions

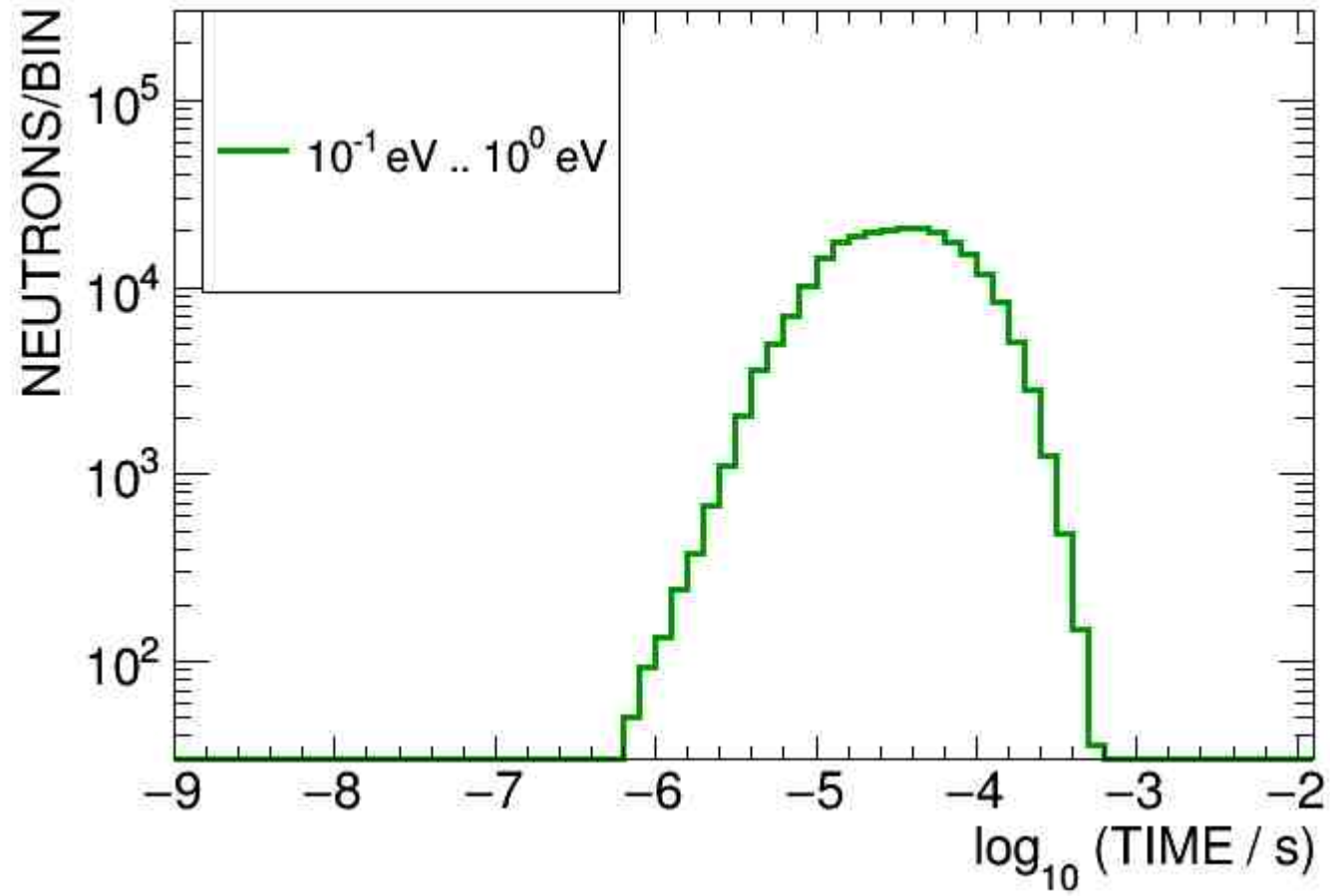


## Proposed setup

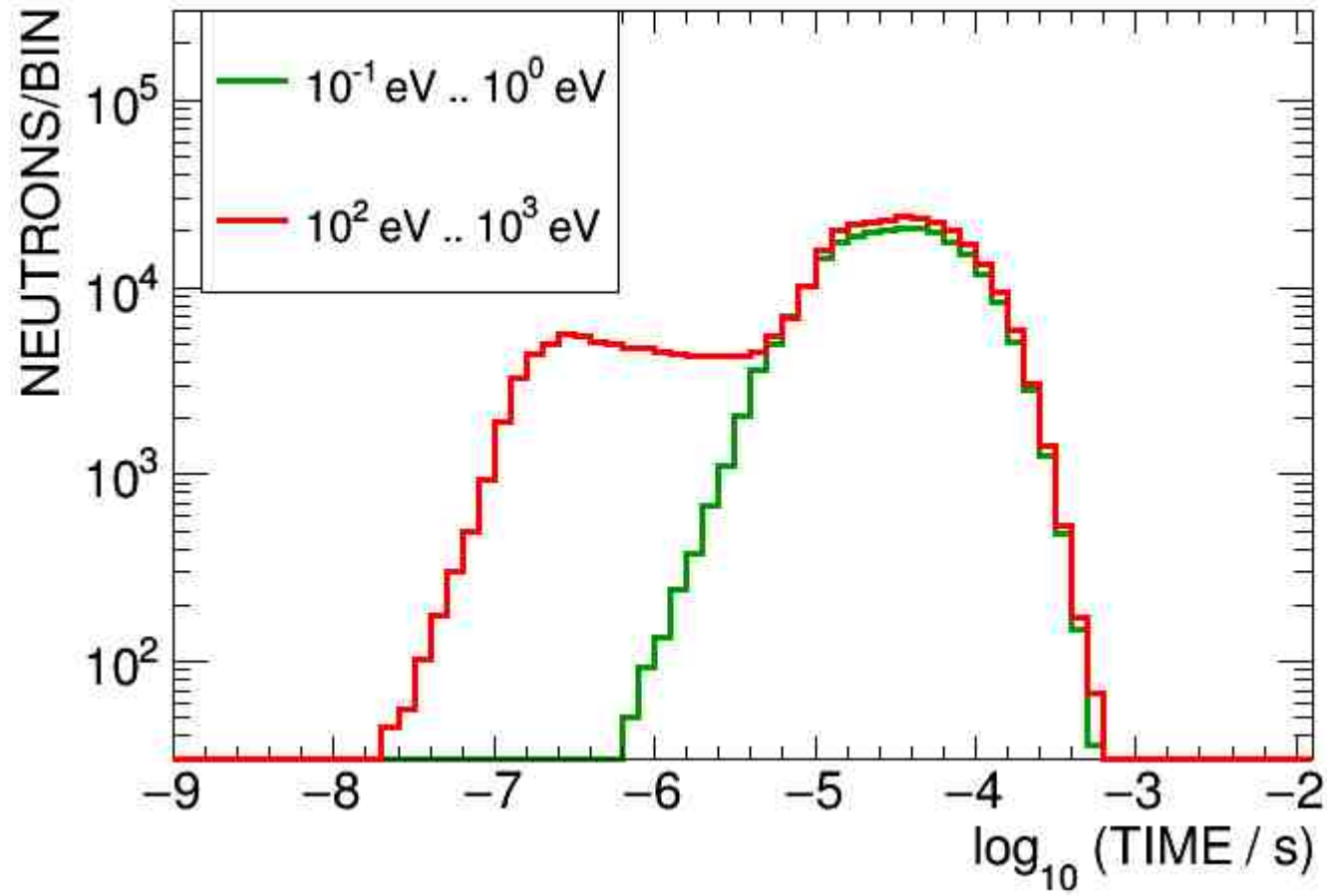
- Moderator: 0.5-2m  $D_2O$
- Spallation target: 10-50 cm W
- Protons: 0.5 – 50 GeV



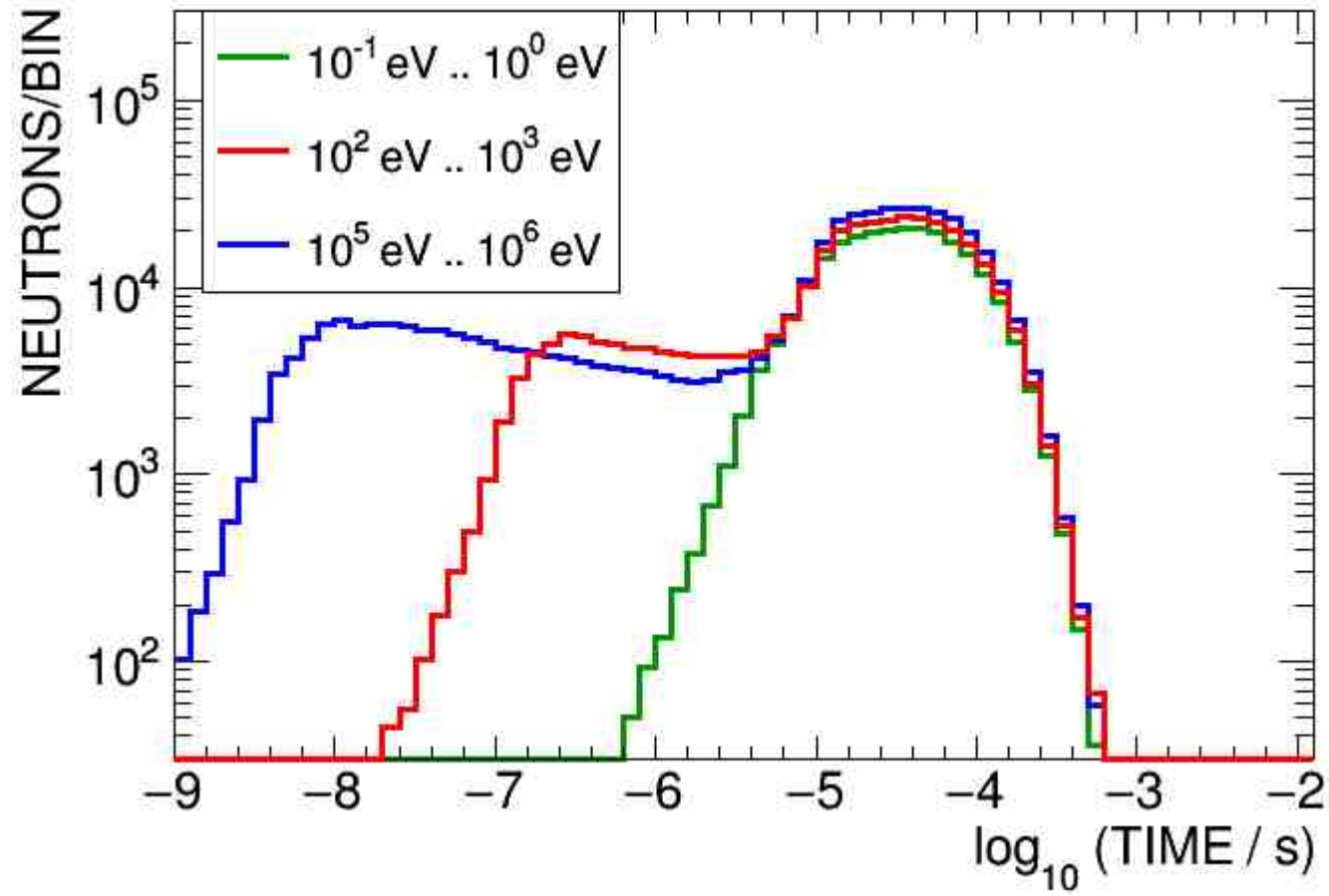
# 1m D<sub>2</sub>O, neutron time in ion pipe



# 1m D<sub>2</sub>O, neutron time in ion pipe

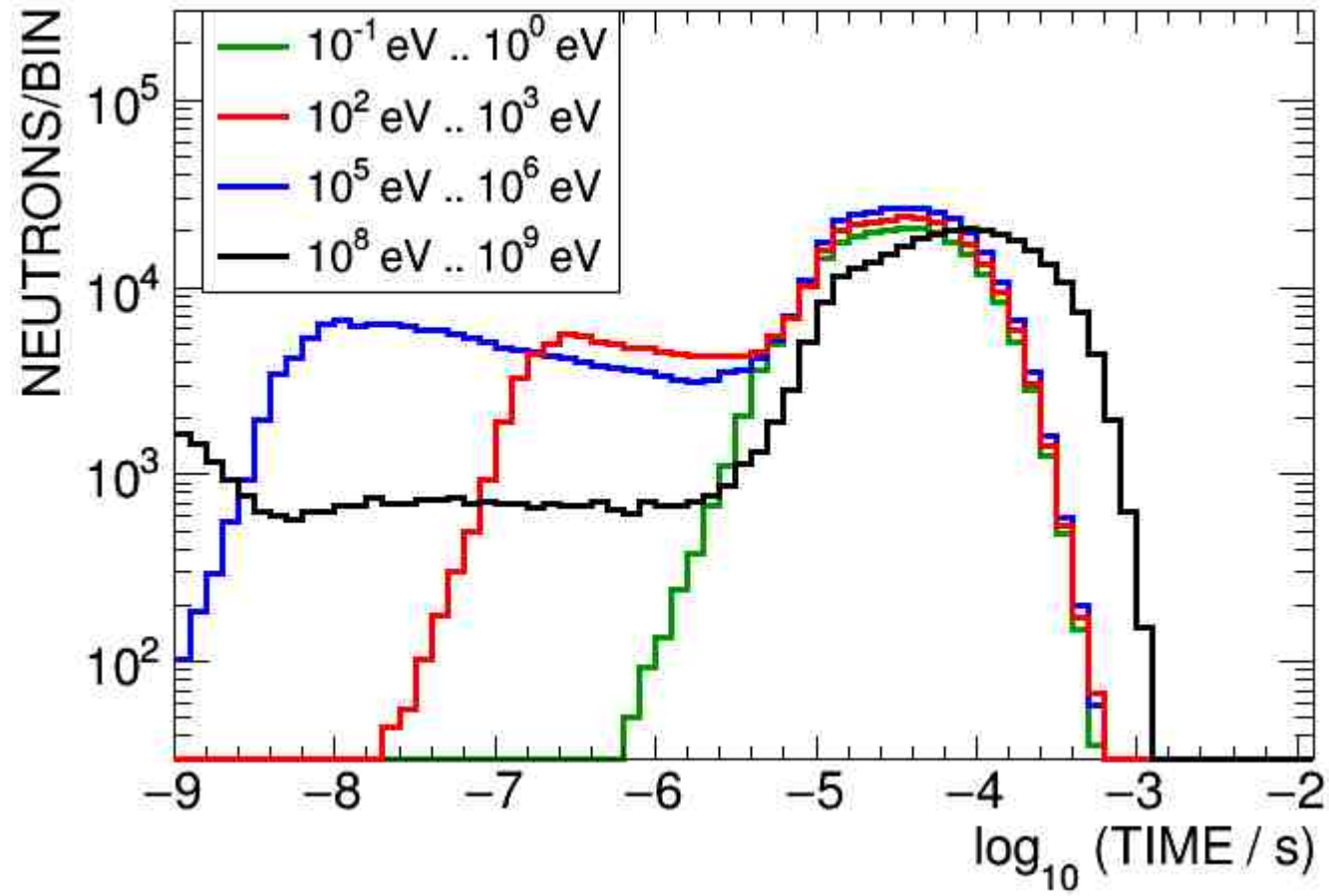


# 1m D<sub>2</sub>O, neutron time in ion pipe

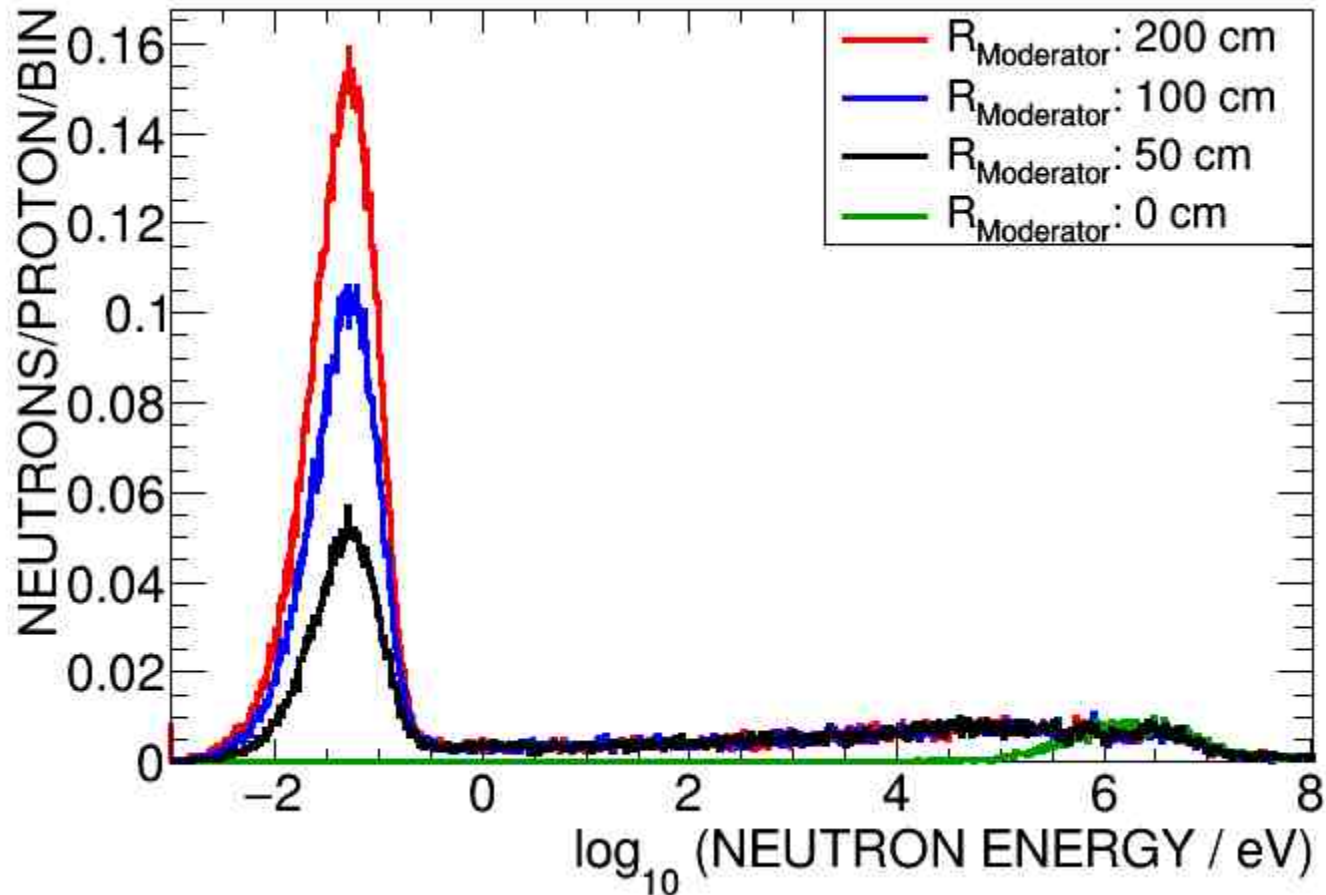




# 1m D<sub>2</sub>O, neutron time in ion pipe



# Energy distribution of target neutrons



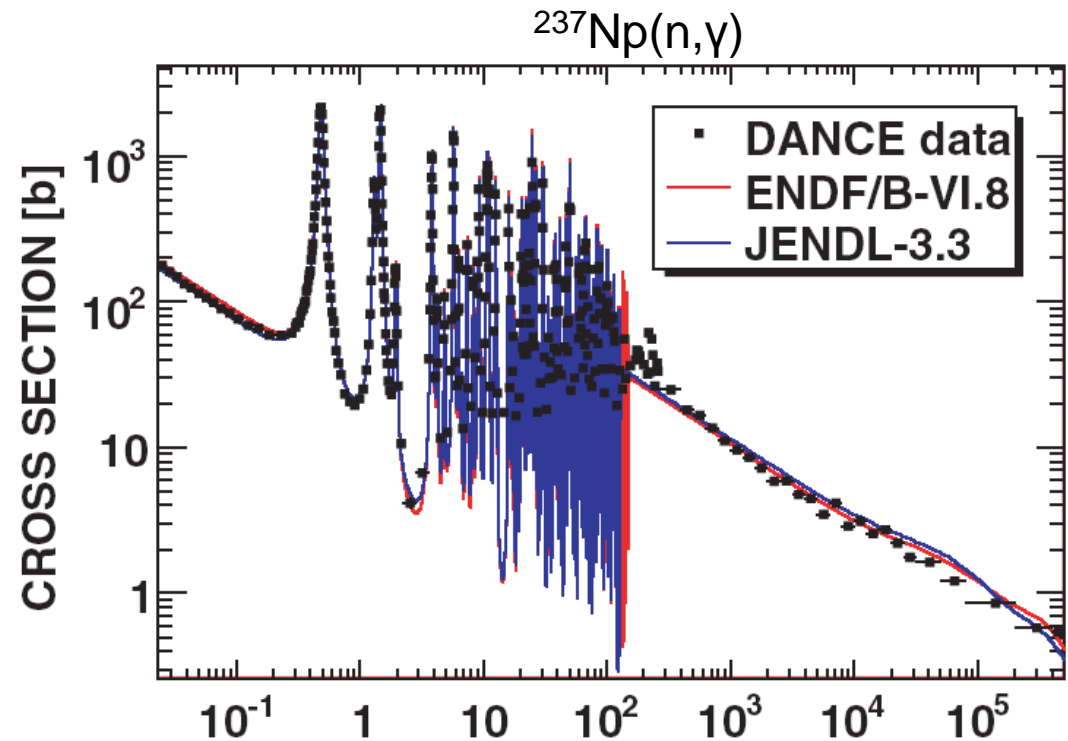
- 800 MeV, 100  $\mu\text{A}$ , 10 cm W, 2 m  $\text{D}_2\text{O}$ :

$$8 \cdot 10^9 \text{ n/cm}^2$$

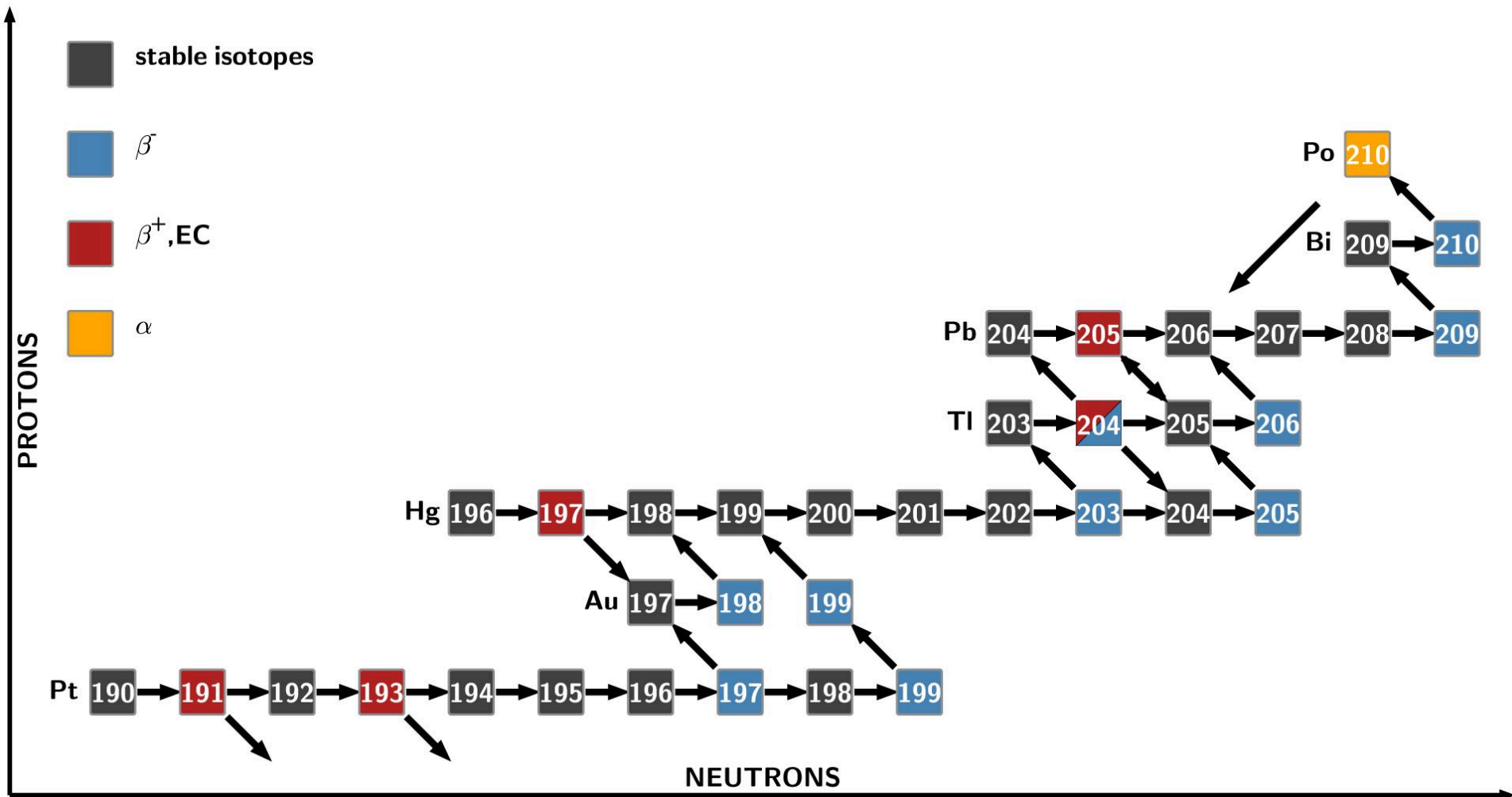
- 20 GeV, 0.5  $\mu\text{A}$  ( $3 \cdot 10^{12}$  protons/s), 50 cm W, 2 m  $\text{D}_2\text{O}$ :

$$5 \cdot 10^8 \text{ n/cm}^2$$

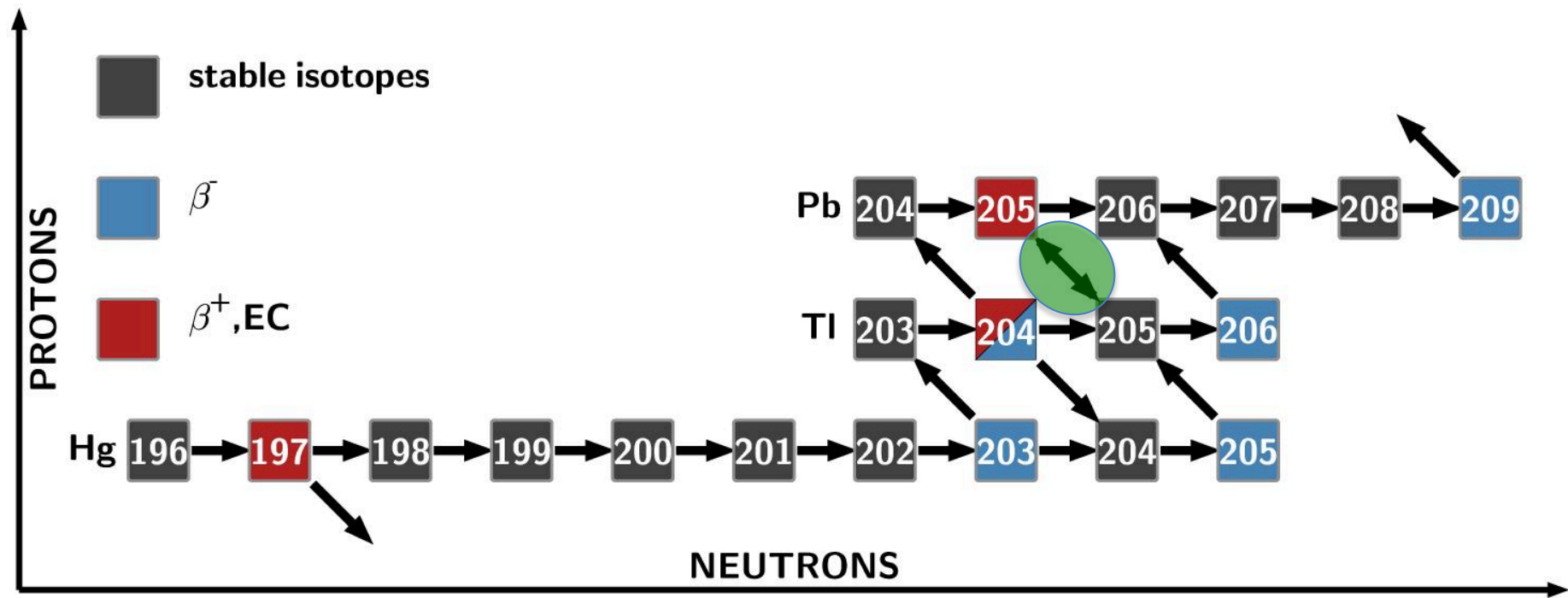
- Protons: 100  $\mu\text{A}$  @ 1 GeV
- **D<sub>2</sub>O moderator radius: 2 m**
- Neutron target:  $10^{10}$  n/cm<sup>2</sup>
- **Counts per day: 10  $\sigma$  / mb**



# The termination point of the s-process



# EC and $\beta^-$ - decay in equilibrium



# Nuclear and atomic physics at work

$^{205}\text{Pb}^{82+} + \text{free } e^{-}$

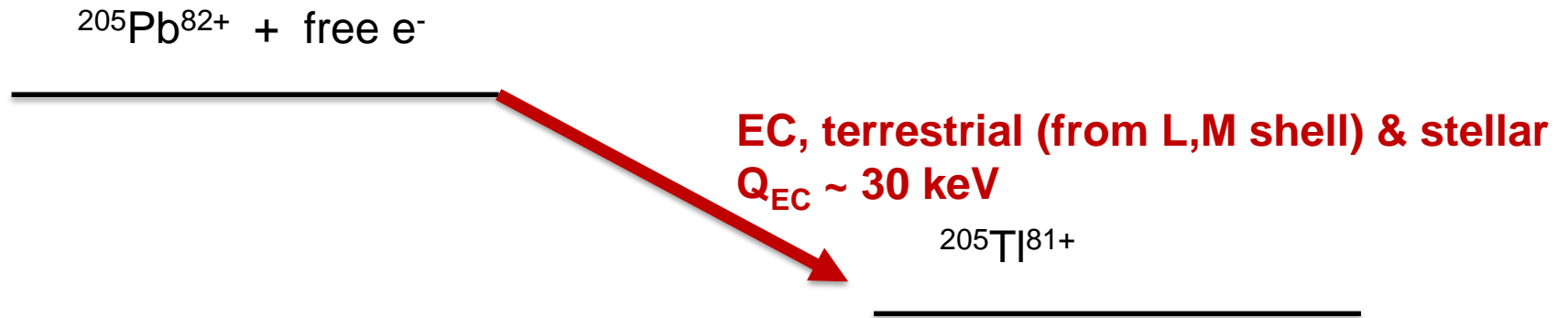
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$^{205}\text{Tl}^{81+}$

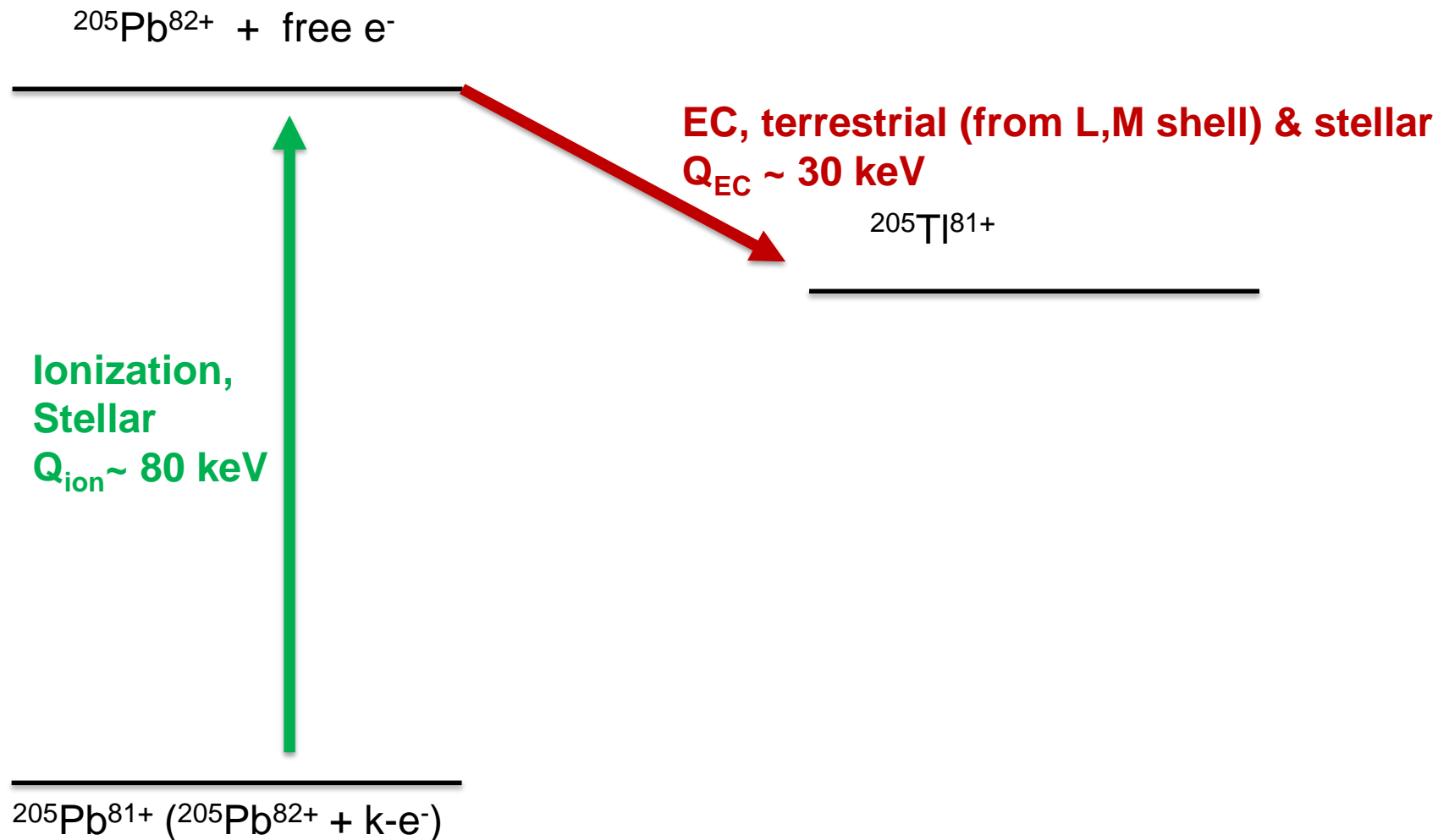
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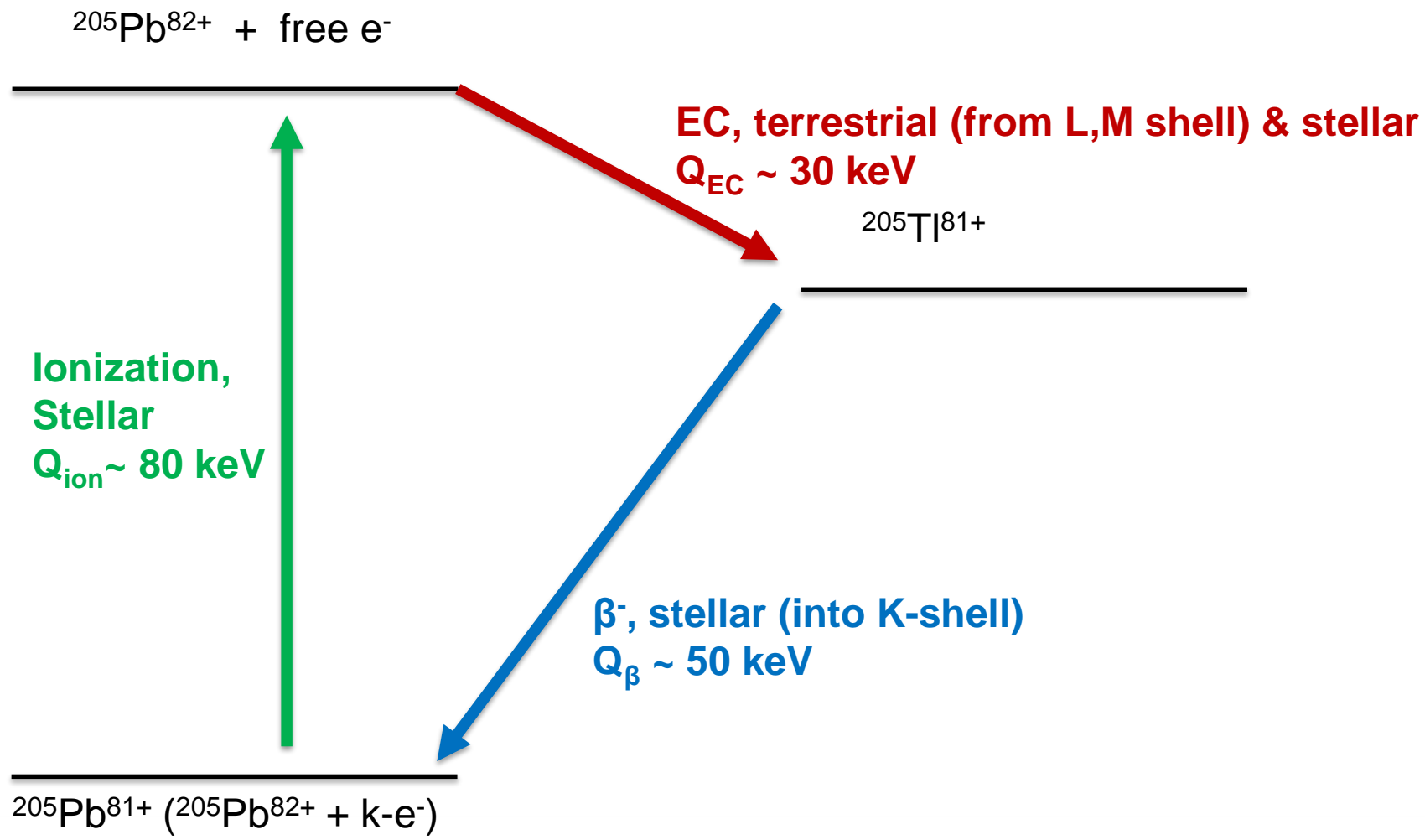
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$^{205}\text{Pb}^{81+} (^{205}\text{Pb}^{82+} + k\text{-}e^{-})$



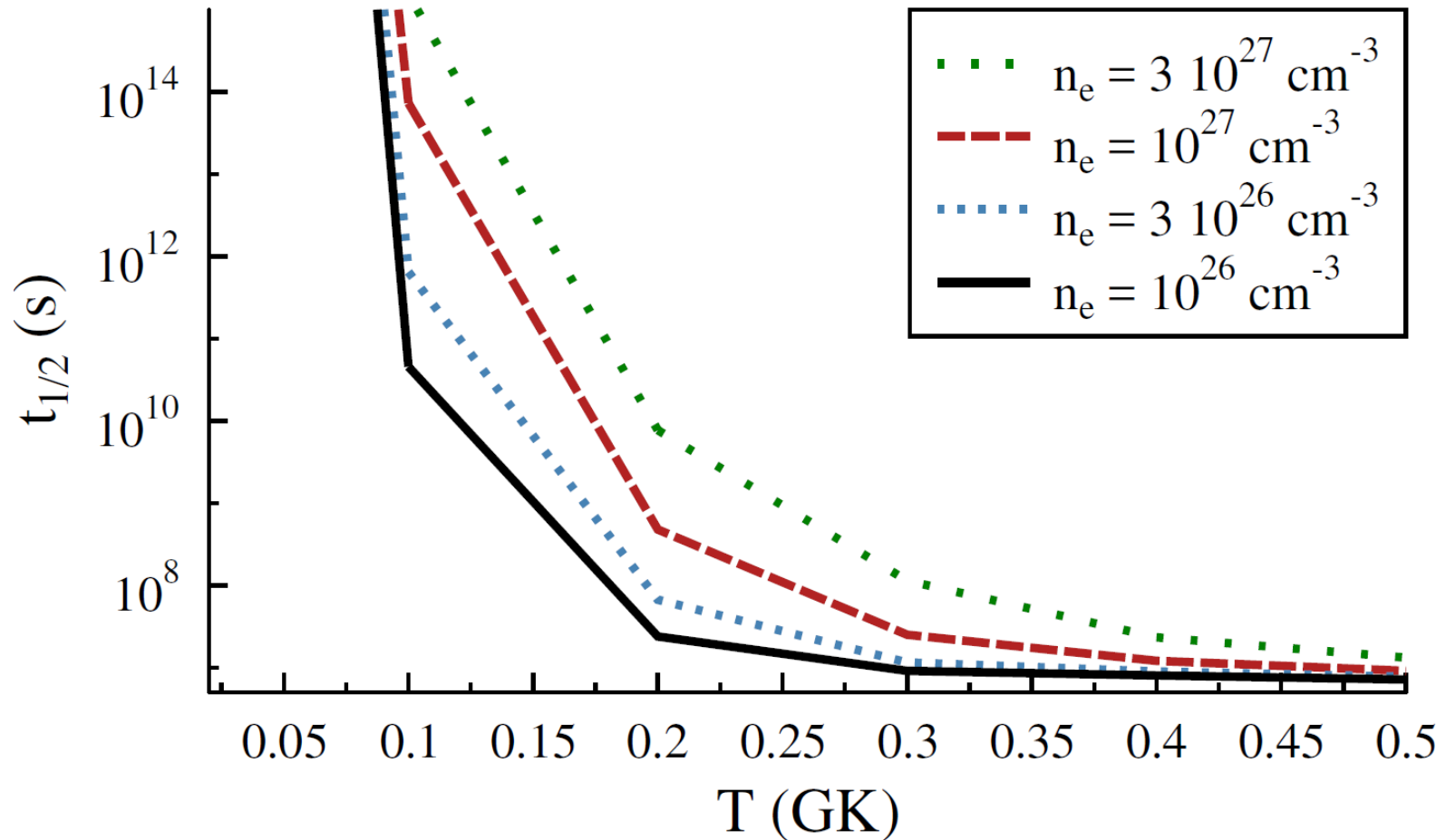






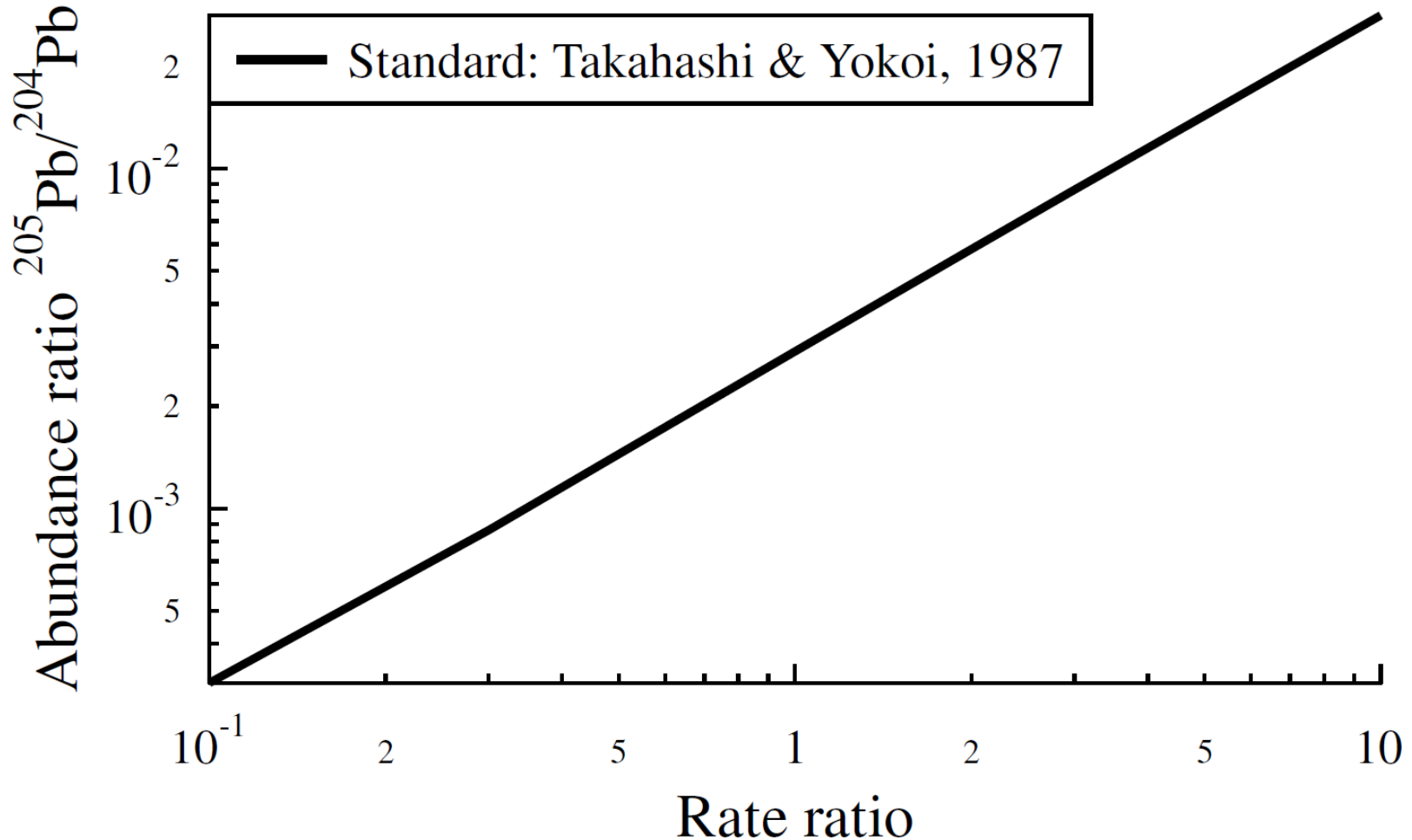
# $\beta^-$ - decay of $^{205}\text{Tl}$

$^{205}\text{Tl}$ , Takahashi & Yokoi, 1987



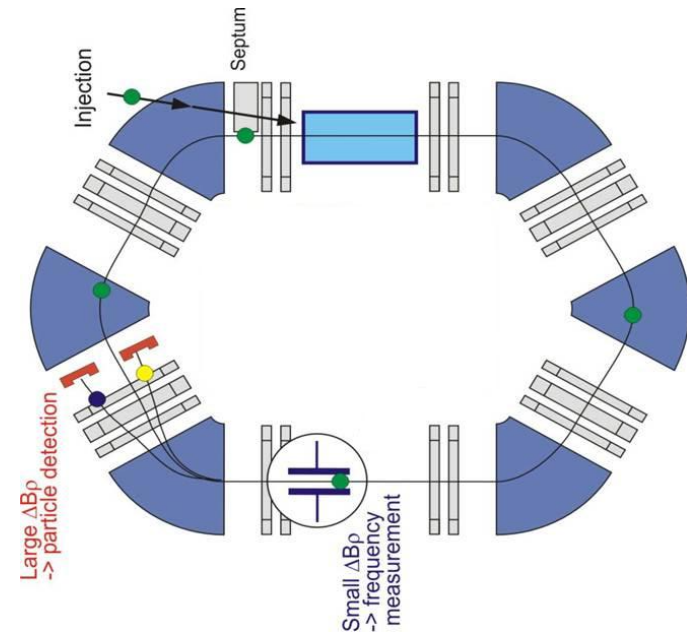
# Equilibrium abundance of $^{205}\text{Pb}$

Impact of  $^{205}\text{Tl}(\beta^-)$  on  $^{205}\text{Pb}$  production



# Measurement of the bound-state beta decay of bare $^{205}\text{Tl}$ ions

- **Proposed at the ESR@GSI/FAIR**
- **Original idea: Fritz Bosch, now Yuri Litvinov**
- **Challenge: Tl is a poison**
  - U-beam
  - Fragmentation
  - Fragment separation
  - Injection into ESR
- **Challenge: detection of decays**
  - Storage, decays:  $^{205}\text{Tl}^{81+} \rightarrow ^{205}\text{Pb}^{81+}$
  - Stripping with Ar-jet:  $^{205}\text{Pb}^{81+} \rightarrow ^{205}\text{Pb}^{82+}$
  - Detection via revolution frequency



# Summary

- Neutron induced reaction studies are very difficult on unstable nuclei
- FRANZ & NAUTILUS will push the limit
- A combination of a neutron target and a ion storage ring might open a new era
- Stellar beta-decay rates are very important to investigate

