# p-process experiments in Cologne

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#### Outline

#### Introduction

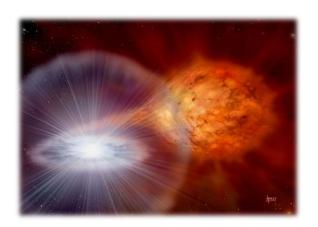
- Nucleosynthesis of heavy elements and the p nuclei
- the  $\gamma$ -process reaction network

# **Experimental measurements of cross-sections**

- $^{112}$ Sn( $\alpha$ , $\gamma$ ) $^{116}$ Te
- $^{108}\text{Cd}(\alpha,\gamma)^{112}\text{Sn}$

The <sup>144</sup>Sm/<sup>146</sup>Sm chronometer



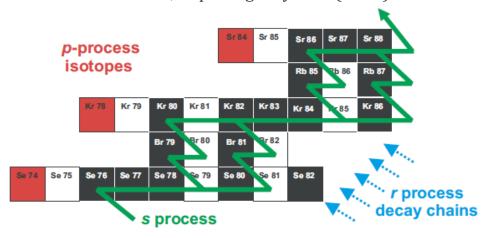


#### The synthesis of the *p* nuclei

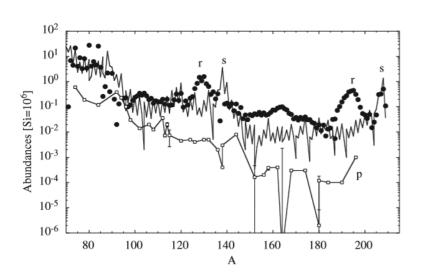
#### p nuclei

- about 30 stable neutron deficient isotopes which cannot be produced by neutron capture reactions
- relatively low isotopic abundances in comparison to s- and r-isotopes
- originally thought to be produced via proton-capture
- temperatures would lead to immediate photodisintegration

T. Rauscher et al., Rep. Prog. Phys. 76 (2013) 066201

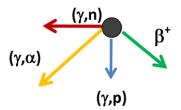


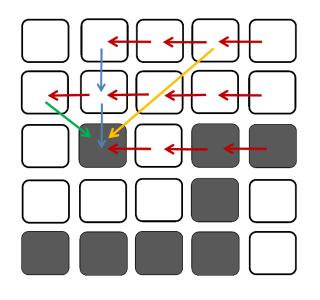
M. Arnould et al. / Physics Reports 450 (2007) 97-213



M. Arnould et al., Phys. Rep. 450 (2007) 97

#### The synthesis of the *p* nuclei



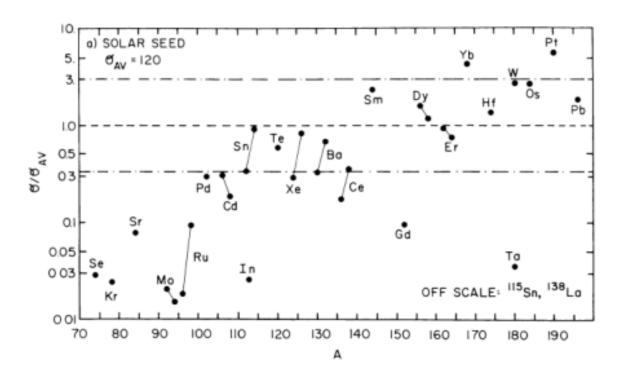


#### γ process reaction-network

- huge photodisintegration reaction-network
- at temperatures between 1.5 GK and 3 GK in ccSN or type Ia SN
- starting from stable seed nuclei formed in the *s*or *r*-process
- $\gamma$ -process path proceeds first via  $(\gamma,n)$  reactions
- branching for A < 130 mainly via (γ,p)</li>
- above A > 130 ( $\gamma$ , $\alpha$ ) get more important

#### The synthesis of the *p* nuclei

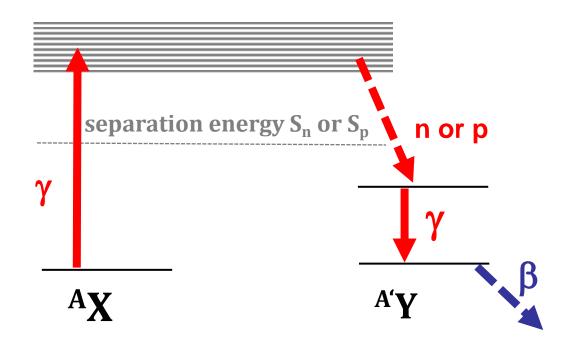
#### γ-process reaction-network calculations



- $\gamma$ -process network calculations cannot reproduce solar system abundance
- other contributions from rp-,  $\alpha$  , vp- other processes?
- problems with reaction rates?

#### Experimental measurements of cross sections

#### **Photodisintegration**



- measuring cross sections via direct detection of ejectiles or via photoactivation
- using either monochromatic  $\gamma$ -ray beams or Bremsstrahlung

#### Experimental measurements of cross sections

#### Statistical model

- cross sections in the Gamow window are small ( $< \mu b$ )
- most of the reactions are not accessible in the laboratory
- reaction rates are calculated mostly in the scope of the statistical model
- cross-section measurements to improve nuclear physics inputparameters:
  - γ-strength function
  - particle + nucleus optical model potentials
  - Nuclear Level Densities

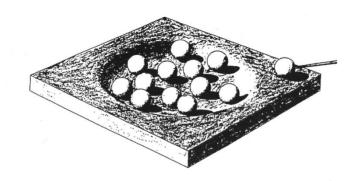
PHYSICAL REVIEW

VOLUME 87, NUMBER 2

IULY 15, 1952

#### The Inelastic Scattering of Neutrons\*

WALTER HAUSER† AND HERMAN FESHBACH
Physics Department and Laboratory of Nuclear Science, Massachusetts Institute of Technology,
Cambridge 39, Massachusetts
(Received March 27, 1952)

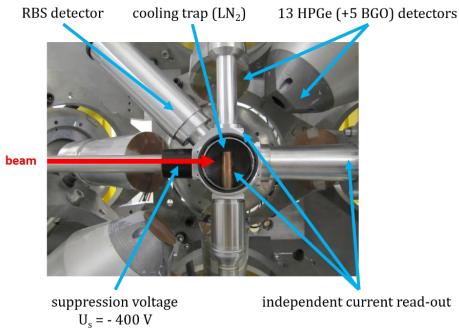


$$\sigma_{jk}^{\mu}(E) = \pi \lambda_{j}^{2} \frac{1}{(2J_{I}^{\mu} + 1)(2J_{j} + 1)} \sum_{J^{\mu}} (2J + 1) \frac{T_{j}^{\mu}(J^{\pi})T_{k}(J^{\pi})}{T_{tot}(J^{\pi})}$$

#### In-Beam γ-ray spectroscopy at HORUS

- 10 MV FN-Tandem accelerator at the IKP Cologne
- HORUS γ-ray spectrometer consists of 13 HPGe detectors (+ RBS)
  - Resolution ≈ 2 keV @ 1332 keV
  - Total efficiency ≈ 2% @ 1332 keV
- Five different angles with respect to the beam axis
  - Determination of angular distributions
- BGO shields for five detectors
- γγ-coincidence measurements



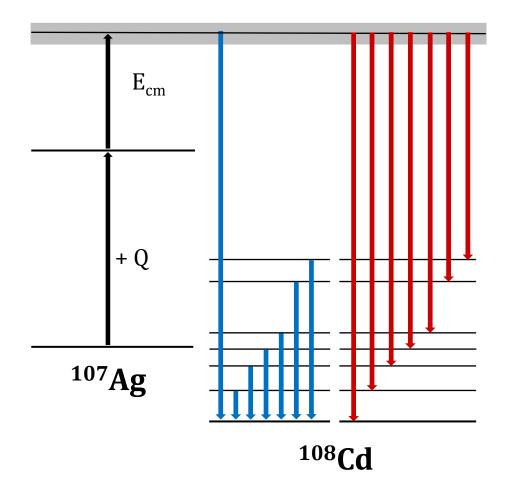


L. Netterdon et al., Nucl. Inst. Meth. A 754 (2014) 94

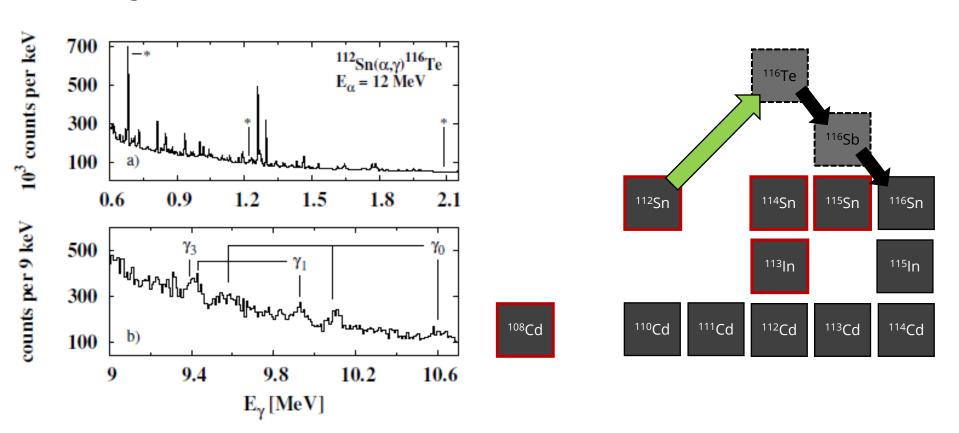
#### In-Beam γ-ray spectroscopy

# In-Beam γ-ray spectroscopy

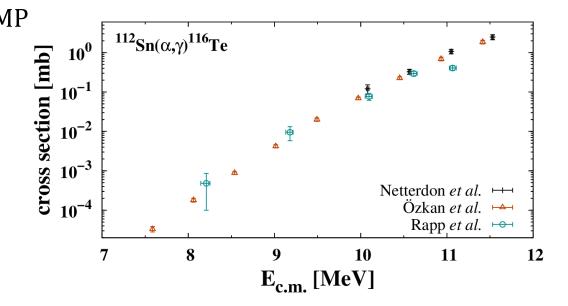
- De-excitation of the entry state
  - Determination of partial cross sections
  - Sensitive to γ-ray strength function
- Transitions to the ground state
  - Determination of the total cross section



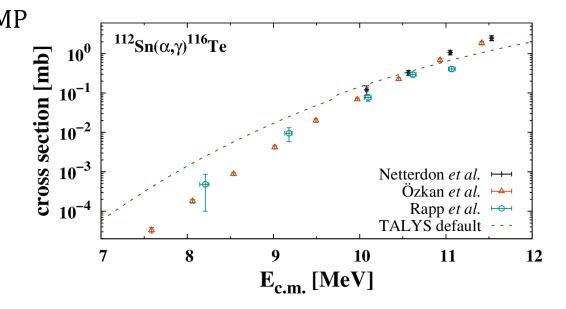
- First in-beam measurement of a radiative  $\alpha$ -capture reaction on a heavy nucleus @ HORUS
- Measurement of total and partial cross sections at four different  $\alpha\text{-}$  energies between 10.5 MeV and 12 MeV



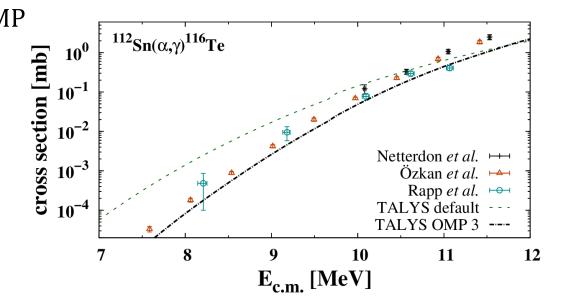
- Measurement confirmed the results of Özkan et al.
- $\bullet$  None of the widely used  $\alpha\textsc{-}\textsc{OMPs}$  led to a good reproduction of the measured cross sections
- Local modifications of the α-OMP could also be used to describe other experimental data of the Cd/Sn- region



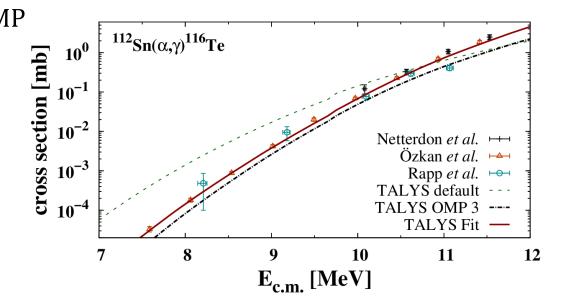
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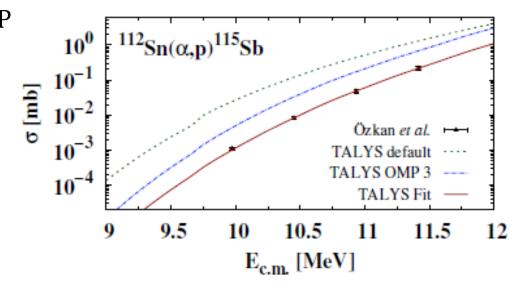


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•  $^{112}$ Sn( $\alpha$ ,p) $^{115}$ Sb

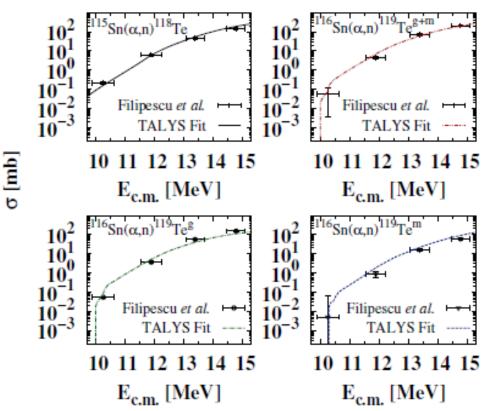


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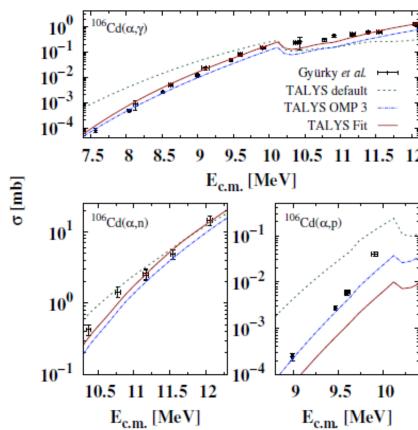
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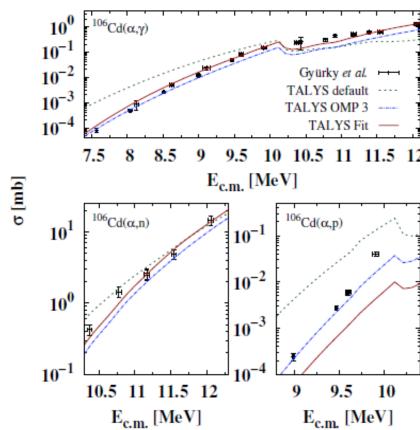
- $^{112}$ Sn( $\alpha$ ,p) $^{115}$ Sb
- $^{115}$ Sn( $\alpha$ ,n),  $^{116}$ Sn( $\alpha$ ,n)



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  - $^{112}$ Sn( $\alpha$ ,p) $^{115}$ Sb
  - $^{115}$ Sn( $\alpha$ ,n),  $^{116}$ Sn( $\alpha$ ,n)
  - $^{106}$ Cd( $\alpha$ ,n),  $^{106}$ Cd( $\alpha$ , $\gamma$ )



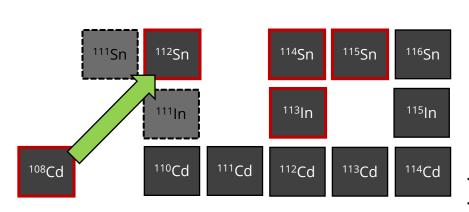
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  - $^{106}$ Cd( $\alpha$ ,n),  $^{106}$ Cd( $\alpha$ , $\gamma$ )



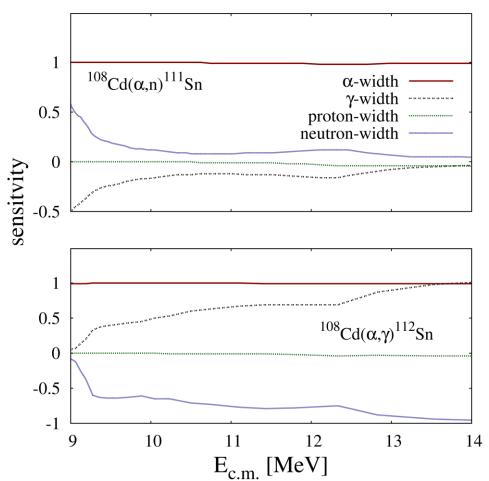
L. Netterdon *et al.*, Phys. Rev. C **91** (2015) 035801

 $\triangleright$  Measurement of  $^{108}$ Cd( $\alpha$ , $\gamma$ ) and  $^{108}$ Cd( $\alpha$ ,n)

# The measurement of $^{108}$ Cd( $\alpha$ , $\gamma$ ) and $^{108}$ Cd( $\alpha$ ,n)

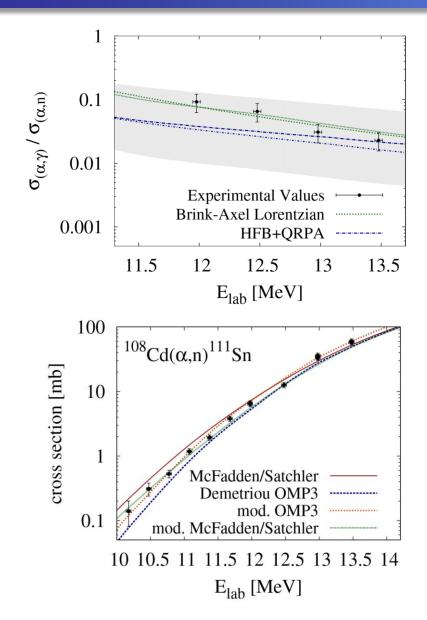


- $^{108}$ Cd( $\alpha$ , $\gamma$ ) very sensitive to other inputparameters
- In-beam Measurement of  $^{108}$ Cd( $\alpha,\gamma$ )
- Activation measurement of  $^{108}$ Cd( $\alpha$ ,n)
- $E_{\alpha}$  between 10.2 MeV and 13.5 MeV
- Target thickness via ICP-MS



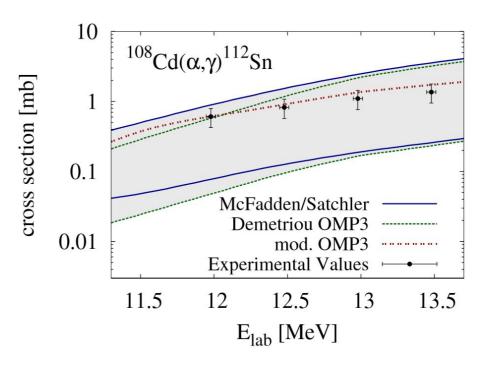
T. Rauscher, The Astrophysical Journal Supplement Series 201, 26 (2012)

# The measurement of $^{108}$ Cd( $\alpha$ , $\gamma$ ) and $^{108}$ Cd( $\alpha$ ,n)



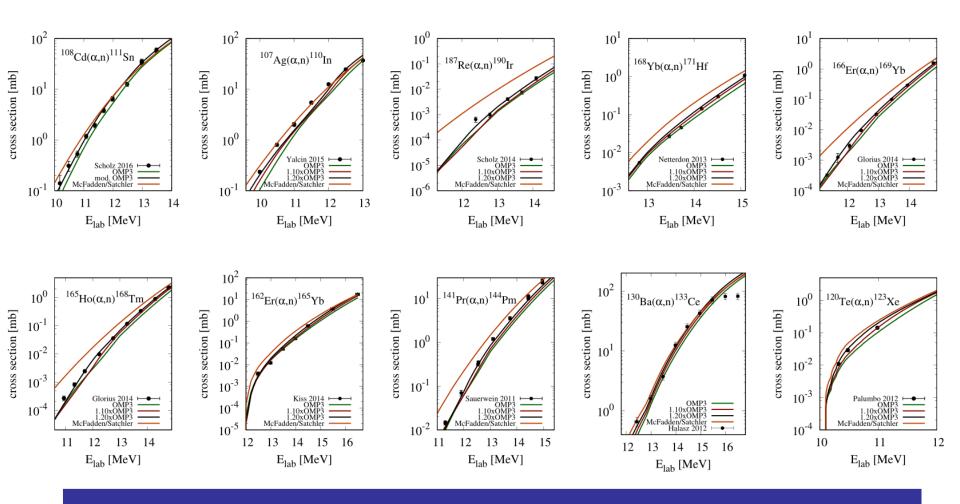
$$\frac{\sigma(\alpha, \gamma)}{\sigma(\alpha, n)} \approx \frac{T_{\alpha}T_{\gamma}}{\sum T_{i}} \times \frac{\sum T_{i}}{T_{\alpha}T_{n}} = \frac{T_{\gamma}}{T_{n}}$$

- Too small values at higher energies
- Wrong energy trend for McFadden
- Best reproduction: modified OMP3



P. Scholz, F. Heim *et al.*, Phys. Lett. B **761** (2016) 247

#### α-optical model potential



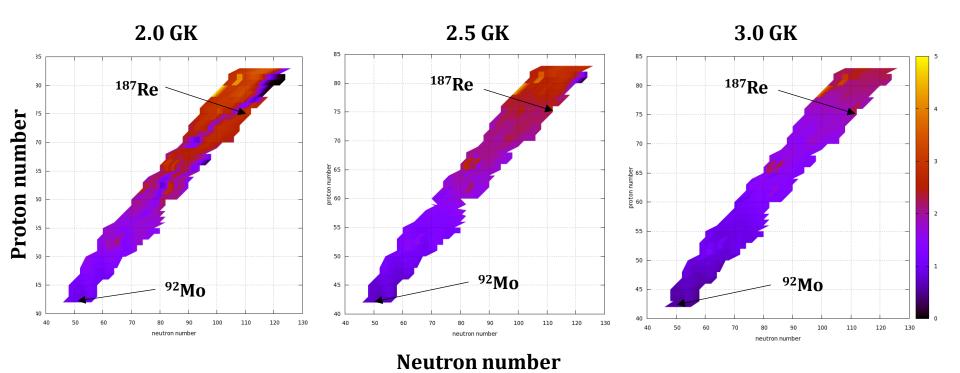
Good reproduction of all experimental  $(\alpha,n)$  data at sub-Coulomb energies

P. Scholz, F. Heim et al., Phys. Lett. B 761 (2016) 247

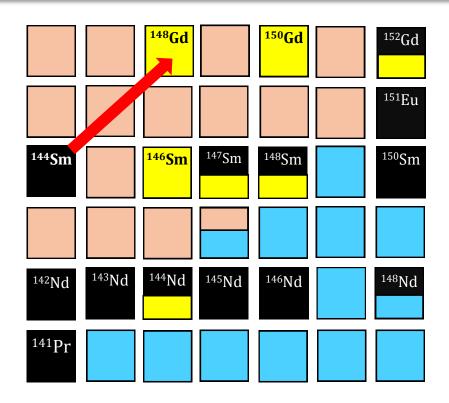
## Comparison McFadden/Satchler and Demetriou OMP3

#### Comparison of $(\gamma,\alpha)$ reaction rates using McFadden/Satchler and Demetriou OMP3

- Factor 10 to 100 difference between A = 100 and A = 140
- 2 to 3 magnitudes difference for A > 140

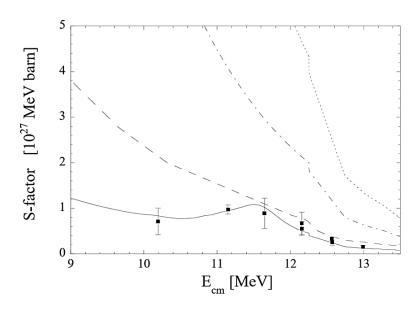


# <sup>144</sup>Sm/<sup>146</sup>Sm chronometer



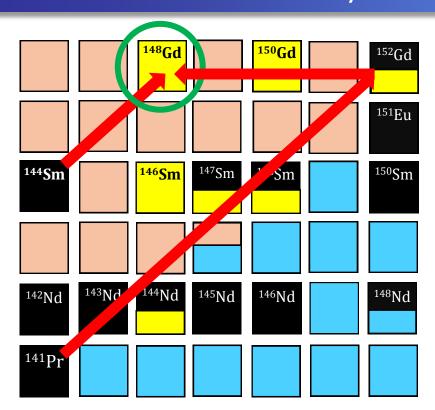
- previous measurement by means of the activation method and α-counting was strongly debated due to methodical reasons
- new measurement via Accelerator Mass Spectrometry (AMS) & counting in Dresden

- extinct p nucleus <sup>146</sup>Sm serves via its α-decay to <sup>142</sup>Nd as a clock for the evolution of our solar system and its planetary processes
- the precise determination of the initial isotopic abundance ratio
   (146Sm/144Sm)<sub>0</sub> is of utmost importance



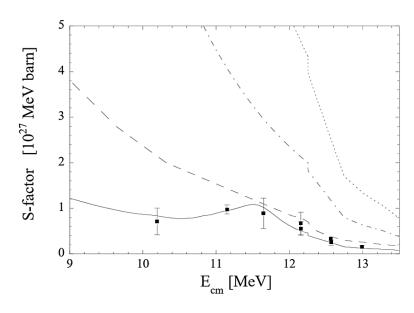
E. Somorjai et al., Astron. Astrophys. 333 (1998) 1112

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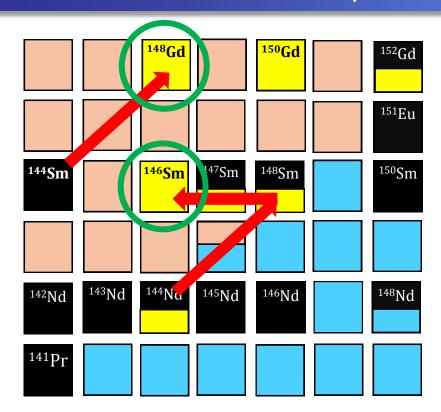
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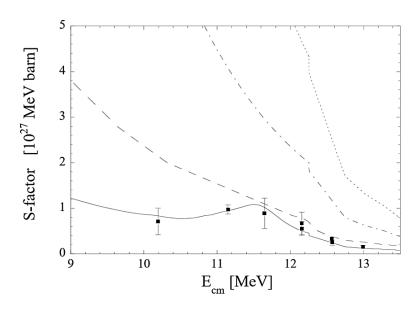
E. Somorjai et al., Astron. Astrophys. 333 (1998) 1112

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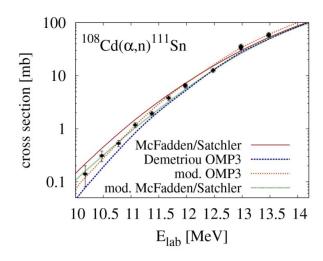
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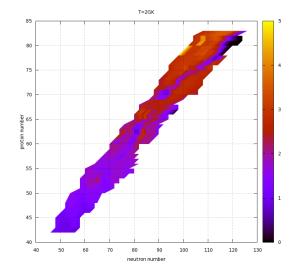


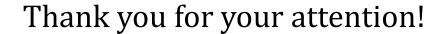
E. Somorjai et al., Astron. Astrophys. **333** (1998) 1112

#### Summary

- Need for robust models for nuclear physics input for a reliable description of nucleosynthesis processes
- α optical-model potential can be used in a wide mass range
- Impact on γ-process has to be studied
- What is about other inputparameters?
- AMS measurements will begin soon in Cologne









R. Altenkirch, V. Derya, A. Dewald, T. Dunai, J. Endres, A. Endres, C. Fransen, F. Heim, S. Heinze, A. Hennig, J. Mayer, C. Müller-Gatermann, C. Münker, L. Netterdon, S.G. Pickstone, S. Prill, M. Schiffer, M. Spieker, P. Sprung, D. Vieß, K.-O. Zell and A. Zilges



J. Meijer and J. Vogt



U. Giesen

H.-W. Becker and D. Rogalla



H. Wilsenach and K. Zuber

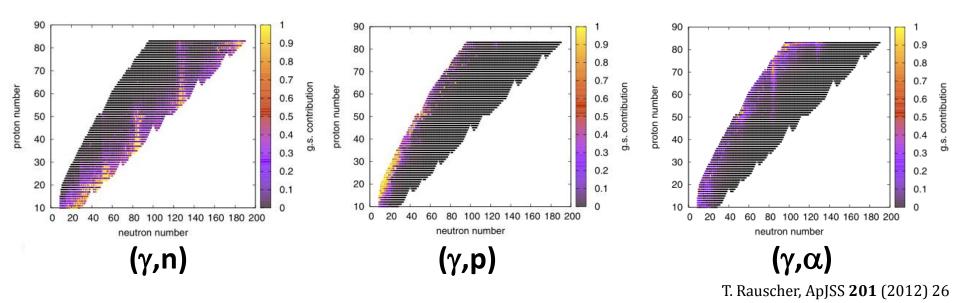


# Backup

#### Experimental measurements of cross sections

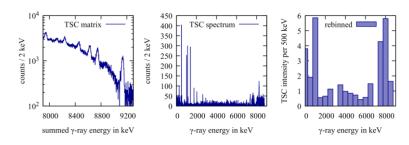
#### **Ground-state contributions**

- measured cross sections cannot directly used for astrophysics
- for  $\gamma$ -induced reaction the ground-state contribution is almost zero
- larger contribution from excited states in the stellar plasma  $(T_9 > 1.5)$
- reaction rates are obtained from the inverse reactions via reciprocity theorem



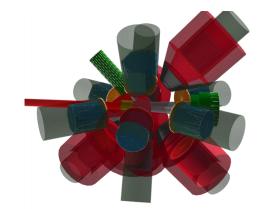
#### Summary

 Need for robust models for nuclear physics input for a reliable description of nucleosynthesis processes



• **Proton-captures** can be used to obtain information about the  $\gamma$ -strength function of even unstable nuclei

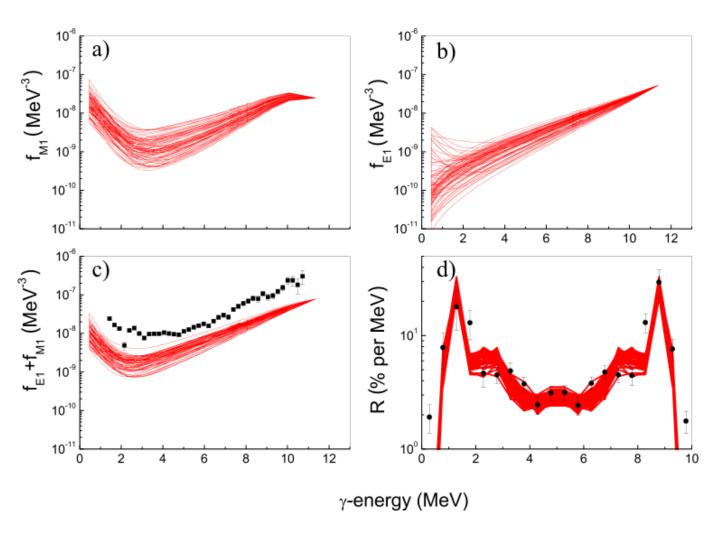
• The method of **Two-Step Cascades** can be used to study  $\gamma$ -strength functions and nuclear level densities





#### Two Step Cascades

#### TSC-analysis for <sup>59</sup>Co(p,γγ) <sup>60</sup>Fe



A. Voinov et al., PRC **81** (2010) 024319

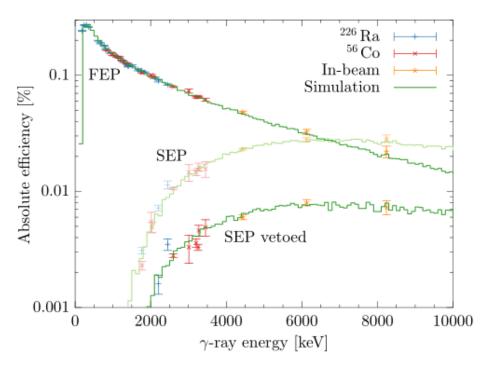
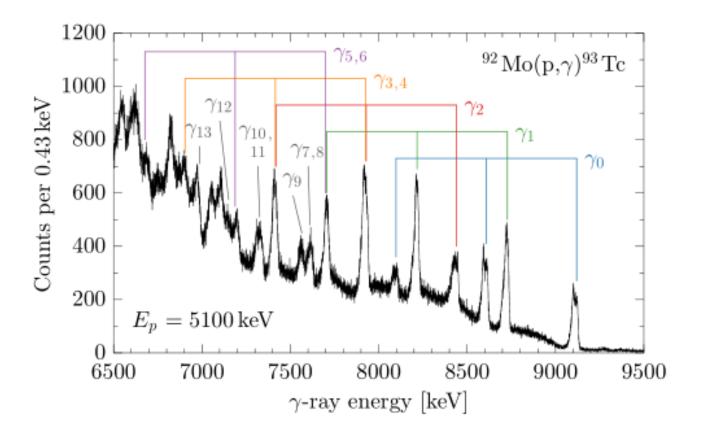


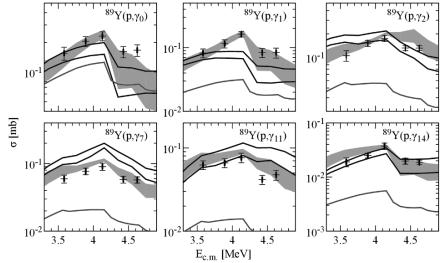
FIG. 2. Full energy peak (FEP) and single escape-peak (SEP) efficiency for one HORUS HPGe detector. Experimental data have been obtained from standard sources and in-beam reactions  $[^{27}\text{Al}(p,\gamma)^{28}\text{Si}, \,^{12}\text{C}(p,p'\gamma), \,^{19}\text{F}(p,\alpha)^{16}\text{O}]$ , with the last two scaled to the FEP efficiency. The simulations were carried out by using GEANT4 [25]. It was not necessary to scale the simulated data to match the experimental results. The detector is equipped with a BGO anti-Compton shield whose active veto signal can be used on an event-by-event basis to suppress escape peaks (SEP vetoed).

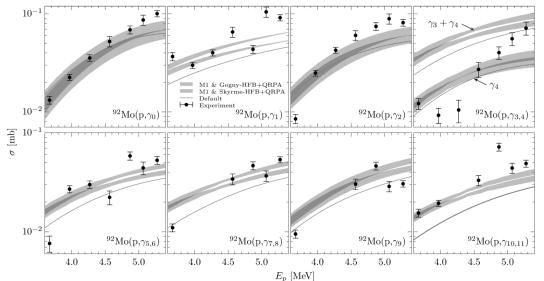


#### In-beam measurement of cross sections

L. Netterdon et al., PLB **744** (2015) 358

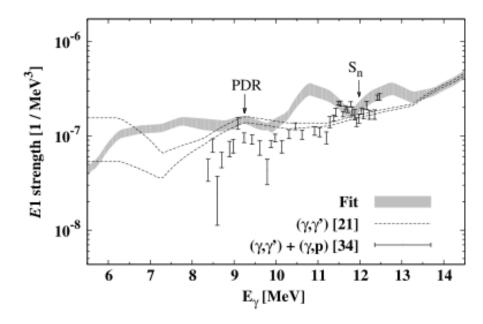
- Measurement of the <sup>89</sup>Y(p,γ)<sup>90</sup>Zr reaction at 5 different proton energies
- Excitation energies up to 13 MeV
- Comparision to  $(\gamma, \gamma')$  data possible





- testing the  $\gamma$ -ray strength function in  $^{93}$ Tc via  $^{92}$ Mo(p, $\gamma$ )
- partial cross sections at 7 different proton energies between 3.5 MeV and 5.5 MeV
- M1/E2-strength not negligble
  - shell model calculations by R. Schwengner for <sup>93</sup>Tc

J. Mayer et al., PRC 93 (2016) 045809



**Fig. 5.** E1-strength distribution as a function of  $\gamma$ -ray energy in  $^{90}$ Zr extracted from the measurement of partial cross sections of the  $^{89}$ Y(p,  $\gamma$ ) $^{90}$ Zr reaction. The gray shaded area depicts the adjusted  $\gamma$ -ray strength function used to reproduce the experimental partial cross sections by Hauser–Feshbach calculations. Additionally shown is the  $\gamma$ -ray strength function obtained from a  $(\gamma, \gamma')$  measurement of Ref. [21] and of Ref. [34]. The dashed lines correspond to the experimental uncertainty of the  $(\gamma, \gamma')$  data. The strength around the PDR energy of about 9 MeV as found in Refs. [11,21] is well reproduced by the present measurement, but a significant enhancement is observed around the neutron-separation energy  $S_n$ .

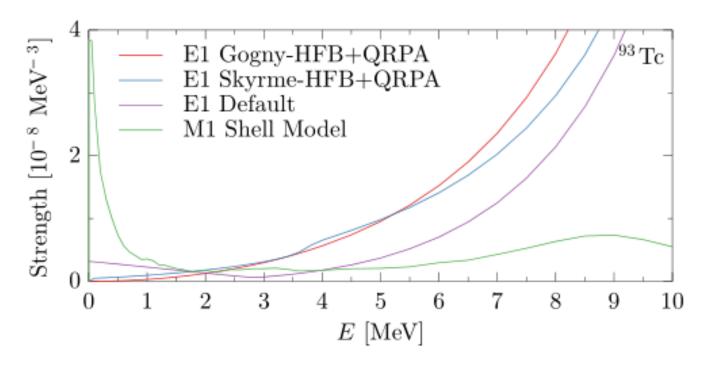
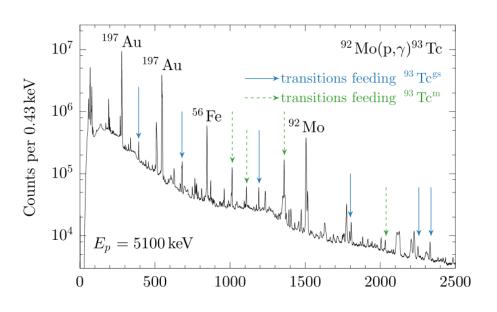
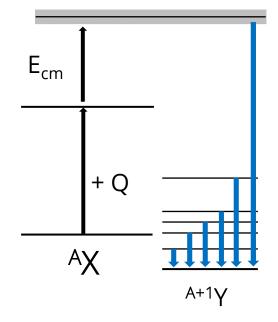


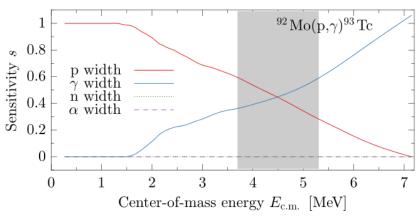
FIG. 10. E1 and M1  $\gamma$  strength in  $^{93}$ Tc obtained by Gogny- or Skyrme-HFB + QRPA in the case of E1 and shell-model calculations for the case of M1 strength. The low-energy M1 strength was combined with a Lorentzian for the spin flip. See text for details.

## Total cross sections of $^{92}$ Mo(p, $\gamma$ ) $^{93}$ Tc

- Measurement of  $^{92}$ Mo(p, $\gamma$ ) $^{93}$ Tc at four energies between 3.6 MeV and 5.3 MeV, thus, above the (p,n) threshold
- Q-Value: 4.1 MeV
- Excitation energies up to 9.2 MeV
- Determination of total yield from angular distribution of γ-rays

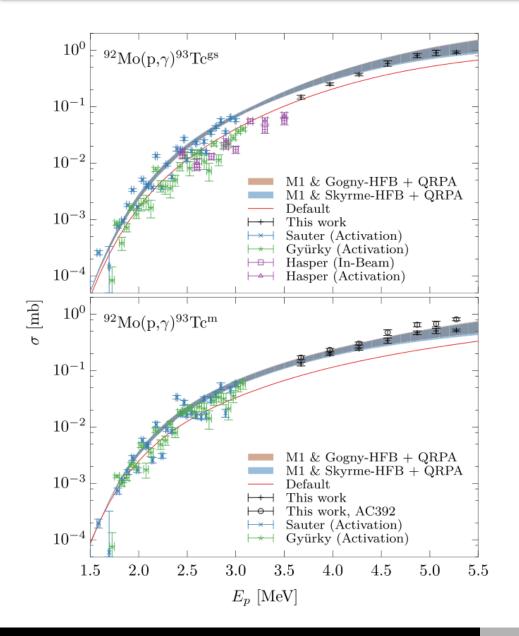






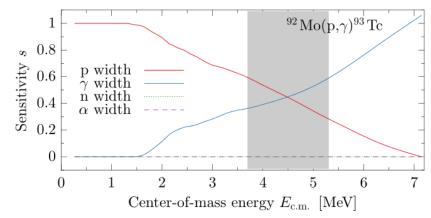
J. Mayer et al., PRC **93** (2016) 045809

## Total cross sections of $^{92}$ Mo(p, $\gamma$ ) $^{93}$ Tc



## Comparison to statistical model calculations with Talys 1.8

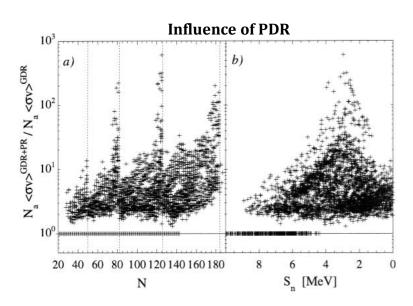
- Measured many times before
- Using different models for γ-strength, NLD, and proton optical model
- Fluctuations at lower energies makes it difficult to draw any conclusions
- Total cross-sections are not enough to find differences in the outcome of various  $\gamma$ -strength models



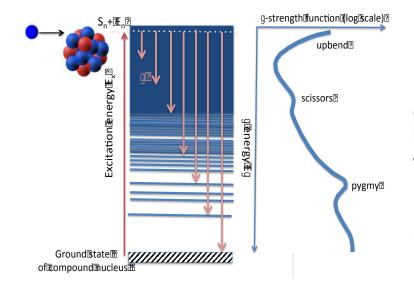
J. Mayer et al., PRC **93** (2016) 045809

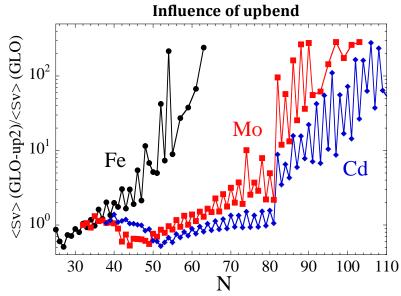
## γ-strength function

- γ-strength plays a crucial role in the determination of reaction rates
- the lower the particle separation energies
  the more important the low energy tail of
  the γ-strength
- Proton-capture can be used to obtain information about the γ-strength function of even unstable nuclei



S. Goriely, Phys. Lett. B 436 (1998) 10

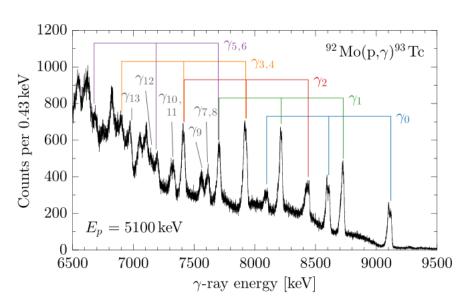


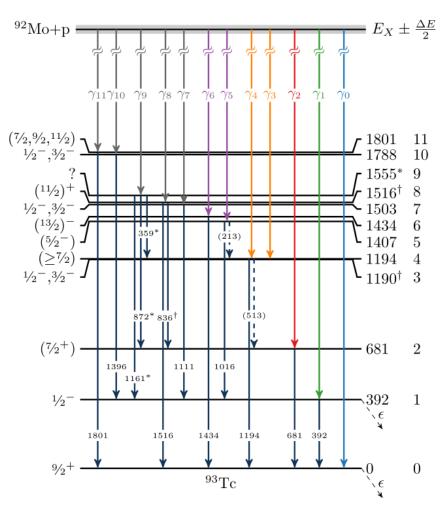


A.C. Larsen and S. Goriely, Phys. Rev. C 82, 014318 (2010)

## Prompt deexcitation and partial cross sections

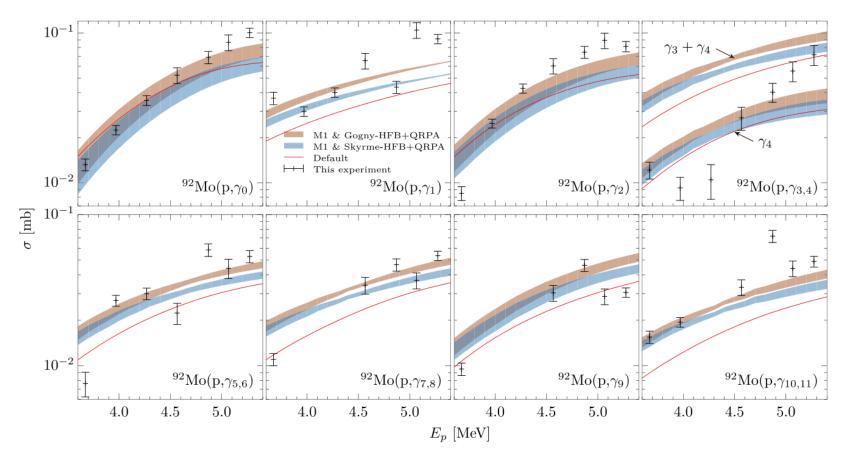
- Detecting the highly energetic prompt  $\gamma$ rays to excited levels in the reaction
  product
- Ratios between partial cross-sections give information about the energy dependence of the γ-strength function
- Adjust γ-strength on partial cross-sections
- Constrain or exclude existing models





J. Mayer et al., PRC **93** (2016) 045809

### In-beam measurement of cross sections



- testing the  $\gamma$ -ray strength function in  $^{93}$ Tc via  $^{92}$ Mo(p, $\gamma$ )
- partial cross sections at 7 different proton energies between 3.5 MeV and 5.5 MeV
- M1/E2-strength not negligble
  - ➤ shell model calculations by R. Schwengner for <sup>93</sup>Tc

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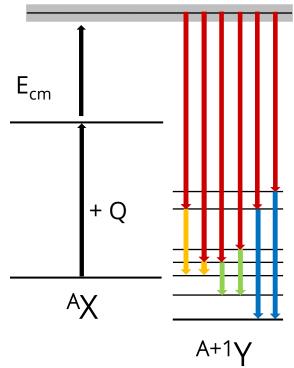
#### Test of photon strength functions by a method of two-step cascades

F. Bečvář and P. Cejnar Charles University, Faculty of Mathematics and Physics, Prague 8, CS-18000, Czechoslovakia

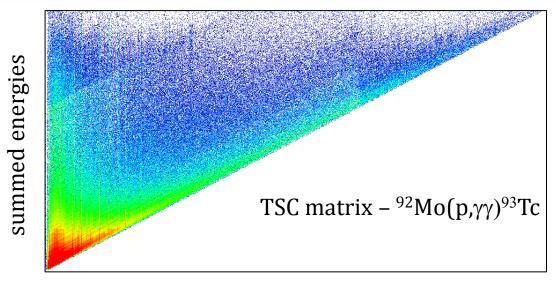
R. E. Chrien
Brookhaven National Laboratory, Upton, New York 11973

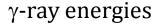
J. Kopecký Nederlands Energy Research Foundation ECN, P.O.Box 1, 1755 ZG Petten, The Netherlands (Received 5 February 1992)

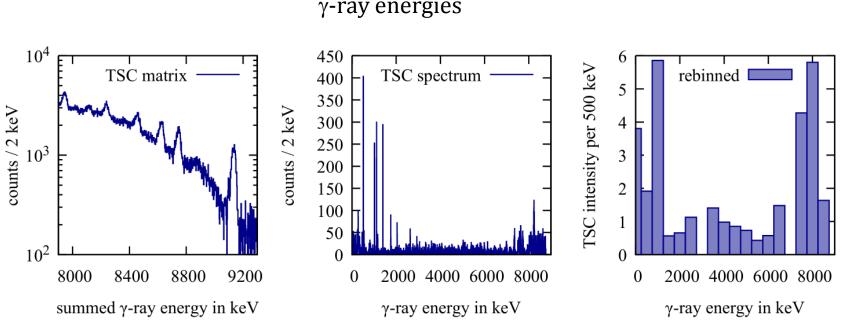
- Detecting two step γ-ray cascades populating states in the reaction product
- TSC spectrum obtained via gate on sum energy of coincident γ-rays
- TSC spectrum can be simulated in the statistical model regarding contribution of E1 or M1 strength
- Different models for nuclear level densities and γ-strength can be tested



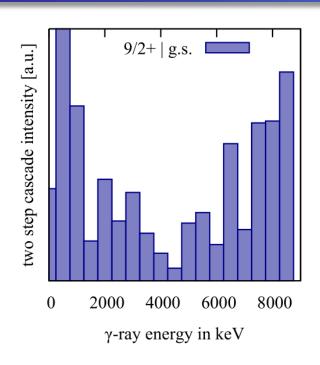
## Two Step Cascades for <sup>92</sup>Mo(p,γγ)<sup>93</sup>Tc

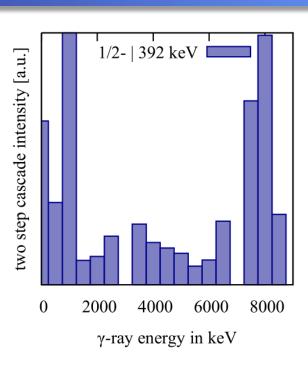


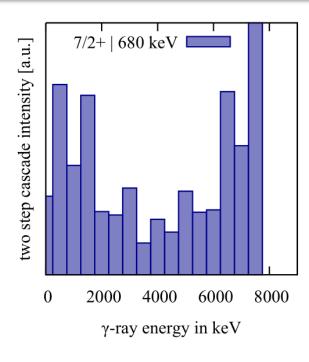




## Two Step Cascades for $^{92}$ Mo(p, $\gamma\gamma$ ) $^{93}$ Tc



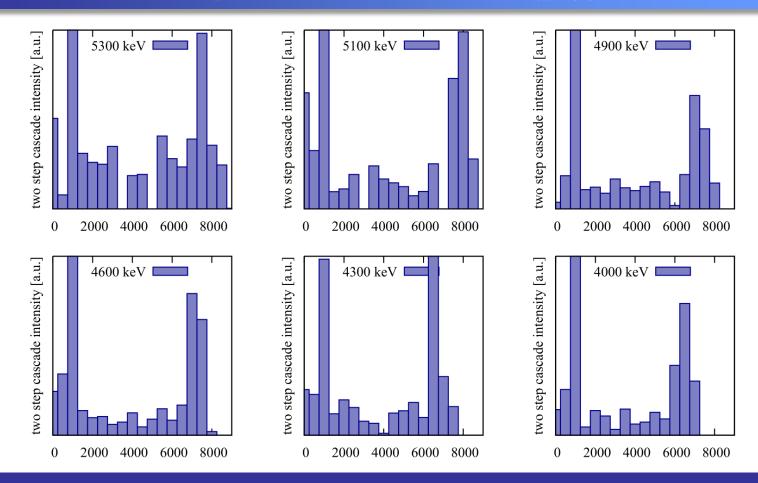




$$I_{if}(E_1,E_2) = \sum_{XL,XL',J_m^\pi} \frac{\Gamma_{im}^{XL}(E_1)}{\Gamma_i} \rho(E_m,J_m^\pi) \; \frac{\Gamma_{mf}^{XL'}(E_2)}{\Gamma_m} \; \; + \sum_{XL,XL',J_{m'}^\pi} \frac{\Gamma_{im'}^{XL}(E_2)}{\Gamma_i} \rho(E_m',J_{m'}^\pi) \; \frac{\Gamma_{m'f}^{XL'}(E_1)}{\Gamma_{m'}} \; . \label{eq:interpolation}$$

Two step cascade intensities to different levels

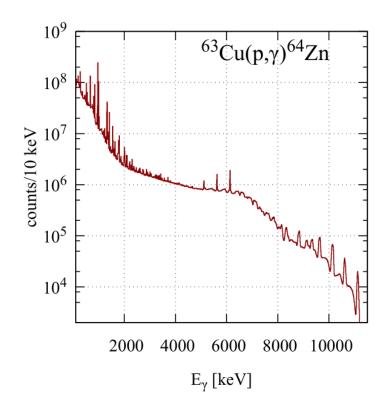
## Two Step Cascades for <sup>92</sup>Mo(p,γγ)<sup>93</sup>Tc



Two step cascade intensities for different excitation energies

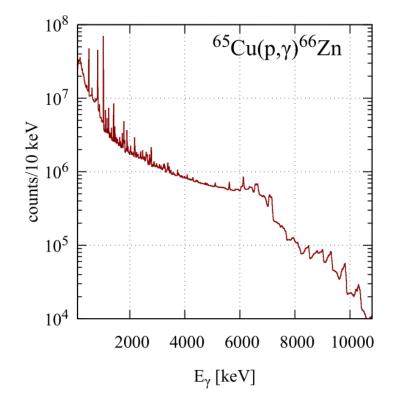
### $^{63}$ Cu(p, $\gamma\gamma$ ) $^{64}$ Zn

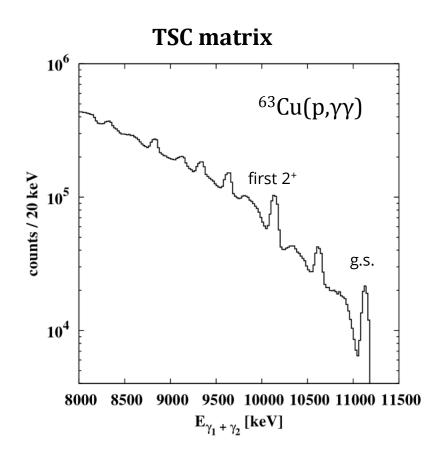
- 4 days with  $\sim$ 400nA
- Target: ~1mg/cm<sup>2</sup>
- Proton energy: 3.5 MeV
- Excitation energies: 11.2 MeV

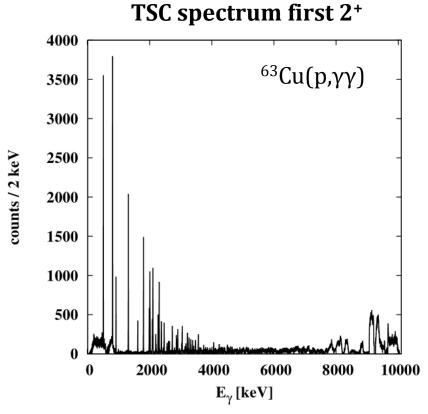


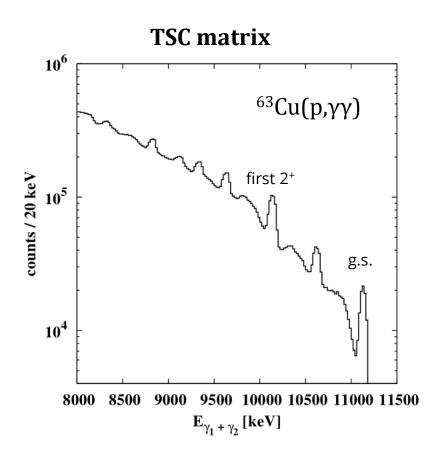
#### <sup>65</sup>Cu(p,γγ) <sup>66</sup>Zn

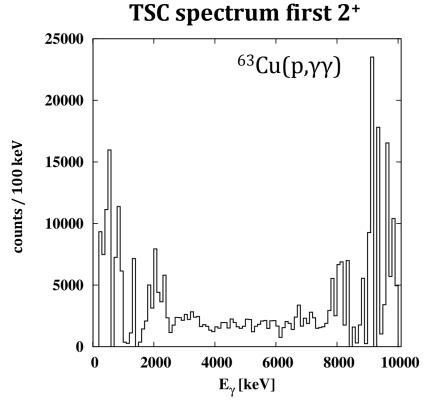
- 5 days with  $\sim$ 500nA
- Target: ~1mg/cm²
- Proton energy: 2.0 MeV
- Excitation energies: 10.9 MeV





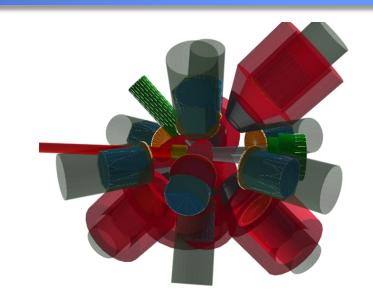




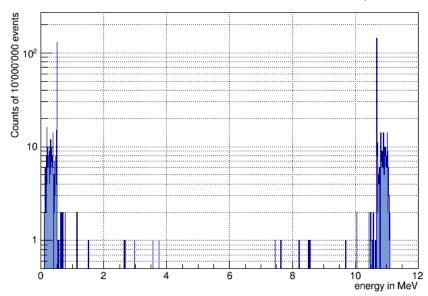


#### **GEANT4 Simulation of the setup**

- Efficiency with  $^{226}$ Ra,  $^{56}$ Co, and  $^{27}$ Al(p, $\gamma$ ) $^{28}$ Si @ 3.6 MeV
- Understanding the response of the detector setup
- Understanding the background in the TSC spectrum



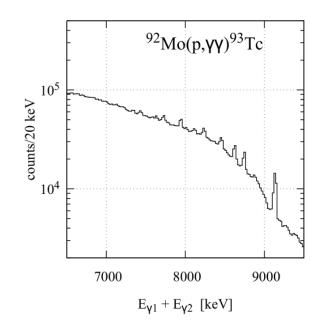
#### Simulated TSC spectrum of single 11.2 MeV γ-ray



#### Outlook

# **Application of TSC analysis on old data**

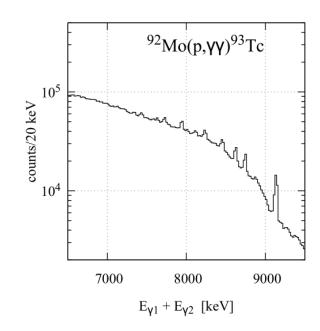
- Coincidence data available for  $(p,\gamma\gamma)$  reactions on  $^{89}$ Y,  $^{92}$ Mo, and  $^{107}$ Ag
- Data available for different beam/excitation energies



#### Outlook

# Application of TSC analysis on old data

- Coincidence data available for  $(p,\gamma\gamma)$  reactions on  $^{89}$ Y,  $^{92}$ Mo, and  $^{107}$ Ag
- Data available for different beam/excitation energies





### **Future experiments**

- New target chamber
- Total and partial cross sections on  $^{109}$ Ag(p, $\gamma$ ) $^{110}$ Cd, Zn, and Ge isotopes
- Other two-step cascade experiments
- Studying of  $(\alpha, \gamma)$  reactions