

MHD assumptions

→ MHD only valid if particles localized:

- by collisions: $\lambda_{mfp} \ll L$ ← plasma scale length
- by B-field: $r_{Larmor} \ll L$

⇒ MHD cannot be used without B-fields!

→ bulk motion of plasma: $V^2 = V_{thermal}^2 + u^2$

particle velocity → V
thermal component due to $f(v)$ → $V_{thermal}$
bulk motion of fluid → u

⇒ macroscopic model

→ slow time scales (ns)

→ the fluid description concept: electrons/ions follow averaged EM fields

- electrons and ions are assumed to form two independent fluids that penetrate each other (2-fluid model)

→ temperatures: T_e & T_i

→ streaming velocities: $\underline{u}_e$ and $\underline{u}_i$

⇒ two shifted distribution functions for the 2 particle populations.
(assume Maxwellian)

$$f_M(v_x) = n \left(\frac{m}{2\pi k_B T} \right)^{1/2} \exp \left(- \frac{m(v_x - u_x)}{2k_B T} \right)$$

(assume ideal gas EOS: $P = \frac{2}{3} U = n k_B T$)

mean drift velocity
in x-direction

Maxwell's equations

→ to derive the MHD equations, must start from the proper combination of Maxwell's equations

→ Gauss law → Poisson eq.: $\underline{E} = -\nabla \Phi$ ← dielectric potential

→ Faraday's Law → EM induction $\overset{\text{B-flux}}{\Phi_m} = \oint_A \underline{B} \cdot d\underline{A} \Rightarrow \underset{\text{ind}}{U} = -\frac{d\Phi}{dt}$

Voltage induced in the loop = -ve change of B-flux in the loop

→ induced voltage depends on:

- B-flux density

- change area A

- changing angle between B-field direction and the area normal ($\underline{B} \cdot d\underline{A}$)

→ Ampere's Law $\oint \underline{B} \cdot d\underline{s} = \mu_0 I$ ← integral form

→ plasmas are diamagnetic: the magnetic moment of a gyreating particle is anti-parallel to the direction of the B-field (mag. moment is a conserved quantity)

⇒ thus we do not use the mag. field strength $\underline{H}$ and dielectric displacement $\underline{D}$, as $\underline{B} = \mu_r \mu_0 \underline{H}$ typical for ferromagnetic materials is not helpful for plasmas as E field is shielded (not dielectric)

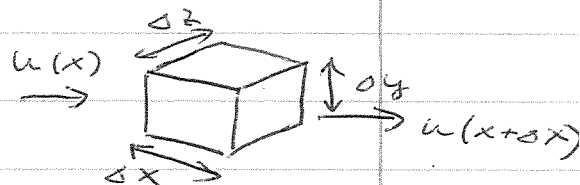
→ in plasmas polarization is a time-varying quantity, static polarization only on plasma surface ⇒ $\underline{D}$ not suitable for static situations

Continuity equation

→ number of particles conserved:

$$N = n \cdot A \cdot \Delta x \quad \text{where area } A = \Delta y \cdot \Delta z$$

→ shown in previous lecture:



→ mass flow through box on LHS in $\Delta t = \rho(x) u(x) \Delta y \Delta z \cdot \Delta t$

→ mass out on RHS in $\Delta t = \rho(x + \Delta x) u(x + \Delta x) \Delta y \Delta z \cdot \Delta t$

⇒ change in mass = $[-\rho(x + \Delta x) u(x + \Delta x) + \rho(x) u(x)] \Delta y \Delta z \Delta t$

⇒ change in $\rho = \frac{\Delta m}{V} = [-\rho(x + \Delta x) u(x + \Delta x) + \rho(x) u(x)] \frac{\Delta t}{\Delta x} = \Delta \rho$

$$\text{as } \Delta x, \Delta t \rightarrow 0 \Rightarrow \frac{\partial \rho}{\partial t} = - \frac{\partial (\rho u)}{\partial x}$$

$$\Rightarrow \text{in 3D: } \boxed{\frac{\partial \rho}{\partial t} + \nabla \cdot (\rho \underline{u}) = 0}$$

→ increase/decrease of particles with changing flow:

$$- \frac{\partial N}{\partial t} = I_N(x + \Delta x) - I_N(x) \approx \frac{\partial I_N}{\partial x} \cdot \Delta x$$

↑ Taylor expand
and retain only the
differential change in flux

→ divide by $\Delta V = A \cdot \Delta x$ and take the limit $\Delta V \rightarrow 0$

$$\Rightarrow \frac{\partial n}{\partial t} + \frac{\partial (n u_x)}{\partial x} = 0 \quad \text{as } I_n = n \cdot \Delta y \cdot \Delta z \cdot u_x$$

$$\Rightarrow \text{in 3D: } \boxed{\frac{\partial n}{\partial t} + \nabla \cdot (n \underline{u}) = 0}$$

Momentum equation

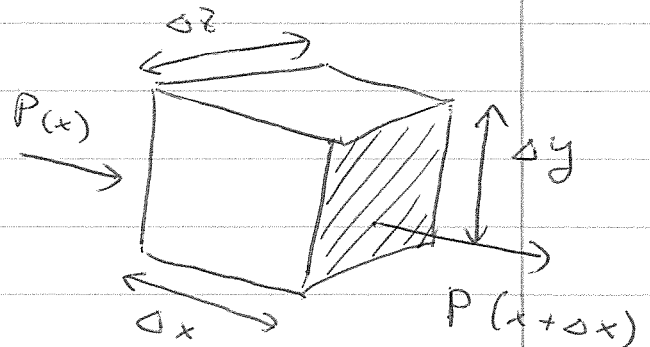
→ net force: sum of all forces acting within the cell

$$F = m \cdot a = m \frac{d\mathbf{v}}{dt} = q(\mathbf{E} + \mathbf{v} \times \mathbf{B})$$

d/dt calculated at the position of point-like particle

→ accelerating fluid:

pressure: $P = \frac{F}{A}$



$$\Rightarrow \Delta F = \Delta P \cdot A, \text{ where } A = \Delta y \Delta z$$

→ in frame moving with the fluid:

$$\Delta F \rightarrow P(x) \Delta y \Delta z - P(x - \Delta x) \Delta y \Delta z = \underbrace{\rho \Delta x \Delta y \Delta z}_{m = \rho \cdot V} \frac{du}{dt}$$

→ then $\Delta x, \Delta y, \Delta z, \Delta t \rightarrow 0$

$$\Rightarrow \text{force per volume: } - \frac{\partial P}{\partial x} = \rho \frac{du}{dt}$$

$$\Rightarrow \text{in 3D: } \boxed{\rho \frac{d\mathbf{u}}{dt} = -\nabla P}$$

$$\hookrightarrow \text{where } \frac{d\mathbf{u}}{dt} = \frac{\partial \mathbf{u}}{\partial t} + (\mathbf{u} \cdot \nabla) \mathbf{u}$$

↑
convective derivative

→ Substitute: $\rho \left[\frac{d\mathbf{u}}{dt} + \mathbf{u} \cdot \nabla \mathbf{u} \right] = -\nabla P$

fixed point
in space

fluid with different velocities
arriving at and leaving at
this point $\Rightarrow$ viscosity

→ number density $\Delta n(v_x)$ of the particle group is related
to the distribution function:

(of velocity between v_x and $v_x + \Delta v_x$)

$$\Delta n(v_x) = \Delta v_x \iint f(v_x, v_y, v_z) dv_y dv_z$$

Note: flow in x-direction only

$\Rightarrow$ momentum flux through a boundary per unit time

→ $|v_x|$ is the rate at which the particles
pass through the boundary

→ net gain/loss: since $I_p = (mv_x) \Delta n(v_x) |v_x| \Delta y \Delta z$

$$\Rightarrow \frac{dP_x}{dt} = -m \frac{d}{dx} (n \langle v_x^2 \rangle \Delta x \Delta y \Delta z)$$

and $P_x - P_{x+\Delta x} = \Delta y \Delta z \cdot m (n \langle v_x^2 \rangle_{x_0-\Delta x} - n \langle v_x^2 \rangle_{x_0})$

$$= \Delta y \Delta z \cdot m (-\Delta x) \frac{d}{dx} (n \langle v_x^2 \rangle)$$

→ note: the sum over Δn results in the average $\langle v_x^2 \rangle$ over the velocity distribution

$$\text{Since } n \langle v_x^2 \rangle = \int f(v_x) v_x^2 dv_x$$

$$\text{and } \Delta n = \Delta v_x \iint f(v_x, v_y, v_z) dv_y dv_z$$

→ by substituting: $v_x = \underset{\substack{\uparrow \\ \text{flow}}}{u_x} + \underset{\substack{\uparrow \\ \text{thermal motion}}}{\tilde{v}_x}$

$$\frac{d}{dt} (nm u_x) = -m \frac{d}{dx} \left[n (u_x^2 + 2 u_x \overset{=0}{\langle \tilde{v}_x \rangle} + \overset{\text{thermal eneg}}{\langle \tilde{v}_x^2 \rangle} \right]$$

$$\frac{1}{2} m \langle \tilde{v}_x^2 \rangle = \frac{1}{2} k_B T$$

$$\langle \tilde{v}_x \rangle = 0, \text{ random motion}$$

$$\text{and } \frac{d(n u^2)}{dx} = u_x \frac{d(n u_x)}{dx} + n u_x \frac{d u_x}{dx}$$

$$\Rightarrow mn \frac{d u_x}{dt} + m u_x \frac{dn}{dt} = -m u_x \frac{d(n u_x)}{dx} - mn u_x \frac{d u_x}{dx} - \frac{d(n k_B T)}{dx}$$

$$\text{but continuity equation: } \frac{dn}{dt} + \frac{d(n u_x)}{dx} = 0$$

$$\Rightarrow \frac{d(n u_x)}{dx} = - \frac{dn}{dt}$$

$$\Rightarrow mn \frac{d u_x}{dt} + \cancel{m u_x \frac{dn}{dt}} = \cancel{+ m u_x \frac{dn}{dt}} - mn u_x \frac{d u_x}{dx} - \frac{d(n k_B T)}{dx}$$

$$\Rightarrow nm \left(\frac{d u_x}{dt} + u_x \frac{d u_x}{dx} \right) = - \frac{d P}{dx} //$$

→ Since we have shown that:

$$\frac{d\underline{u}}{dt} = \frac{\partial \underline{u}}{\partial t} + \underline{u} \cdot \nabla \underline{u}$$

MHD = HD + Lorentz force

$$\Rightarrow \underline{F}_L = -e \underbrace{(\underline{v}_e \times \underline{B})}_{\text{electrons}} + Ze \underbrace{(\underline{v}_i \times \underline{B})}_{\text{ions}}$$

and we have current density: $\underline{j} = nq\underline{v}$

$$\underline{F} = q\underline{v} \times \underline{B} \rightarrow \underline{F} \cdot \underline{h} = nq \underbrace{\underline{v} \times \underline{B}}_{\underline{j} \times \underline{B}}$$

$$\Rightarrow \frac{\underline{F}}{V} = \underline{F} \cdot \underline{h} = \underline{j} \times \underline{B} = \nabla P$$

• two fluid model:

$$n_e m_e \left[\frac{\partial \underline{u}_e}{\partial t} + (\underline{u}_e \cdot \nabla) \underline{u}_e \right] = -n_e e (\underline{E} + \underline{u}_e \times \underline{B}) - \nabla P_e$$

$$n_i m_i \left[\frac{\partial \underline{u}_i}{\partial t} + (\underline{u}_i \cdot \nabla) \underline{u}_i \right] = +n_i e (\underline{E} + \underline{u}_i \times \underline{B}) - \nabla P_i$$

→ space charge: $\rho = n_i e - n_e e$

→ current density: $\underline{j} = n_i e \underline{u}_i - n_e e \underline{u}_e$

→ and since $\frac{F_L}{V} = \underline{j} \times \underline{B} = \nabla P$ (force per volume)

⇒ momentum equation for MHD:

$$\rho \left[\frac{d\underline{u}}{dt} + \underline{u} \cdot \nabla \underline{u} \right] = \underline{j} \times \underline{B} - \nabla P$$

→ can add other forces

Moments of the distribution function

→ moment = weighted average (averaged over the distribution function)

$$n(\underline{r}) = \int f(\underline{r}, \underline{v}) d^3v \Rightarrow \text{number density}$$

→ bulk motion velocity $\underline{u}(\underline{r})$ integral of $f(\underline{v})$ over all $\underline{v}$'s

$$\Rightarrow \text{weighted average of } \underline{v} = \int \underline{v} f(\underline{r}, \underline{v}) d^3v = \underline{u}(\underline{r}) n(\underline{r})$$

$$\therefore \underline{u}(\underline{r}) = \frac{\int \underline{v} f(\underline{r}, \underline{v}) d^3v}{\int f(\underline{r}, \underline{v}) d^3v}$$

$$\rightarrow 0^{\text{th}} \text{ order moment: } n(\underline{r}) \propto \int \underline{v}^0 f d^3v$$

$$\rightarrow 1^{\text{st}} \text{ order moment: } \underline{u}(\underline{r}) \propto \int \underline{v} f d^3v$$

$$\rightarrow 2^{\text{nd}} \text{ order moment: } \underline{U}(\underline{r}) \propto \int \underline{v} \cdot \underline{v} f d^3v$$

↑
thermal energy

} obtain P and T through the equation of state

Generalized Ohm's Law

start from momentum equations for electrons and ions!

ions: $n m_i \frac{d\mathbf{u}_i}{dt} = n e (\mathbf{E} + \mathbf{u}_i \times \mathbf{B}) - \nabla P_i + n m_i \mathbf{g} + n \nu_{ei} m_e (\mathbf{u}_e - \mathbf{u}_i)$ $\nwarrow \tilde{\nu}_{pe i}$
①

electrons: $n m_e \frac{d\mathbf{u}_e}{dt} = -n e (\mathbf{E} + \mathbf{u}_e \times \mathbf{B}) - \nabla P_e + n m_e \mathbf{g} + n \nu_{ei} m_e (\mathbf{u}_i - \mathbf{u}_e)$ ②

$\nearrow$ gravity $\nearrow$ collisions

neglecting the $\mathbf{u} \cdot \nabla \mathbf{u}$ term for slowly evolving plasma

with momentum exchange between electrons and ions
being described by the collision frequency ν_{ei}

→ the mean mass motion:

$$\rho_m \frac{d\mathbf{v}_m}{dt} = \mathbf{j} \times \mathbf{B} - \nabla p + \rho_m \mathbf{g}$$

$\nwarrow$ total current density → Lorentz force

where: $\rho_m = n(m_i + m_e)$

(static situation)
when $\frac{d\mathbf{v}_m}{dt} = 0$

→ and total pressure: $p = p_e + p_i$

→ mean mass velocity: $\mathbf{v}_m = \left(\frac{m_i \mathbf{u}_i + m_e \mathbf{u}_e}{m_e + m_i} \right)$

$\nwarrow$ reduced mass

→ in the Lorentz force $\mathbf{j} \times \mathbf{B}$ acts on the total current density. There is no effect of friction between electrons and ions (fluids) as it does not change the total momentum (only influences redistribution).

→ multiply eq. (1) by m_e and (2) by m_i & subtract:

$$\Rightarrow n m_i m_e \frac{d}{dt} (\underline{u}_i - \underline{u}_e) =$$

$$n e (m_e + m_i) \underline{E} + n e (m_e \underline{u}_i + m_i \underline{u}_e) \times \underline{B} - m_e \nabla P_i + m_i \nabla P_e + n (m_e + m_i) v_{ei} m_e (\underline{u}_e - \underline{u}_i)$$

→ neglect m_e in the sum of masses → very small

⇒ the mixed term:

$$\begin{aligned} m_e \underline{u}_i + m_i \underline{u}_e &= m_i \underline{u}_i + m_e \underline{u}_e \\ &\quad + m_i (\underline{u}_e - \underline{u}_i) + m_e (\underline{u}_i - \underline{u}_e) \\ &= \frac{1}{n} \rho_m \underline{v}_m - (m_i - m_e) \frac{1}{n e} \underline{j} \end{aligned}$$

where we define $\underline{j} = (\underline{u}_i - \underline{u}_e) n e$ and $\rho_{e,i} = n \cdot m_{e,i}$

$$\Rightarrow \frac{m_i m_e}{e} \cdot \frac{d \underline{j}}{dt} = e \rho_m (\underline{E} + \underline{v}_m \times \underline{B} - \frac{v_{ei} m_e}{n e^2} \underline{j})$$

$$- m_i \underline{j} \times \underline{B} - \cancel{m_e \nabla P_i} + m_i \nabla P_e \quad \left. \begin{array}{l} \text{divide by} \\ n \end{array} \right\} \begin{array}{l} \text{Hall effect} \\ \text{thermal pressure} \end{array}$$

small

→ slowly varying phenomena $\Rightarrow \frac{d \underline{j}}{dt} = 0$, neglect $\frac{m_e}{m_i}$ terms

⇒ generalized Ohm's law:

$$\underline{E} + \underline{v}_m \times \underline{B} = \eta \underline{j} + \frac{1}{n e} (\underline{j} \times \underline{B} - \nabla P_e)$$

where plasma resistivity: $\eta = \frac{v_{ei} m_e}{n e^2}$ (Coulomb collisions)

Induction equation

- electron/ion velocity components: 1) random thermal motion
2) bulk average motion
3) motion in flow of current $\underline{j}$

→ start from the simplified Ohm's Law:
(ignore the Hall and thermal pressure terms)

$$\underline{E} + \underline{u} \times \underline{B} = \eta \underline{j} = \frac{\eta}{\mu_0} \nabla \times \underline{B}$$

and substitute Ampere's Law: $\nabla \times \underline{B} = \mu_0 \underline{j}$

$$\Rightarrow \underline{E} = \frac{\eta}{\mu_0} \nabla \times \underline{B} - \underline{u} \times \underline{B}$$

then using the Faraday's Law:

$$\frac{\partial \underline{B}}{\partial t} = -\nabla \times \underline{E} = -\nabla \times \left(\frac{\eta}{\mu_0} \nabla \times \underline{B} \right) + \nabla \times (\underline{u} \times \underline{B})$$

$$\therefore \frac{\partial \underline{B}}{\partial t} = -\nabla \times \underline{E} = -\frac{\eta}{\mu_0} \nabla \times (\nabla \times \underline{B}) + \nabla \times (\underline{u} \times \underline{B}) //$$

Magnetic diffusion

→ induction equation in a static case: $\underline{v} = 0$:

$$+ \frac{\partial \underline{B}}{\partial t} + \frac{\eta}{\mu_0} \nabla \times (\nabla \times \underline{B}) = + \frac{\partial \underline{B}}{\partial t} + D_B \nabla^2 \underline{B} = 0$$

$\nabla \times \nabla \times$ (with arrow pointing to $\nabla \times (\nabla \times \underline{B})$)

where diffusion coefficient $D_B = \frac{\eta}{\mu_0}$

describes the diffusion of B-field in a conducting medium

→ estimate diffusion time: assume exponential decay of the field $\underline{B}(t) \propto \exp(-t/\tau_B)$

$$\Rightarrow -\frac{\partial B}{\partial t} = D_B \cdot \Delta B \Rightarrow \frac{\partial B}{\partial t} = -\frac{1}{\tau_B} \cdot B$$

diffusion time scale (with arrow pointing to τ_B)

for $\underline{B}(t) = B_0 \cdot \exp(-t/\tau_B)$

$$\Rightarrow \frac{B}{\tau_B} = D_B \Delta B \quad \text{and} \quad \Delta B \approx \frac{B}{l^2} \quad (\Delta = \nabla^2 \sim \frac{1}{l^2})$$

where l is the characteristic plasma length

$$\Rightarrow \frac{B}{\tau_B} = D_B \frac{B}{l^2} \Rightarrow \frac{1}{\tau_B} = \frac{\eta}{\mu_0 l^2}$$

$$\therefore \tau_B = \frac{\mu_0 l^2}{\eta} //$$

Frozen-in Magnetic Flux

• for $\eta \rightarrow 0$, $R_m \rightarrow \infty$, induction reaction becomes:

$$\frac{\partial \underline{B}}{\partial t} = \nabla \times (\underline{v} \times \underline{B})$$

• continuity equation: $\frac{\partial n}{\partial t} + \nabla \cdot (n \underline{u}) = 0$

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \underline{j} = 0$$

$$\Rightarrow -\frac{1}{\rho} \frac{d\rho}{dt} = \frac{1}{\rho} \nabla \cdot (\rho \underline{v})$$

$$\Rightarrow \frac{d\rho}{dt} + \rho \cdot \nabla \cdot \underline{v} = 0$$

$$\Rightarrow \nabla \cdot \underline{v} = -\frac{1}{\rho} \cdot \frac{d\rho}{dt}$$

$$\text{as } \frac{d\underline{u}}{dt} = \frac{d\underline{u}}{dt} + \underline{u} \cdot \nabla \underline{u} \quad (\text{shown earlier})$$

→ vector identity: $\nabla \times (\underline{A} \times \underline{B}) = (\underline{B} \cdot \nabla) \underline{A} + \underline{A} (\nabla \cdot \underline{B}) - (\underline{A} \cdot \nabla) \underline{B} - \underline{B} (\nabla \cdot \underline{A})$

⇒ apply the vector identity to the above expression:

$$\begin{aligned} \nabla \times (\underline{v} \times \underline{B}) &= (\underline{B} \cdot \nabla) \underline{v} - (\underline{v} \cdot \nabla) \underline{B} + \underbrace{\underline{v} (\nabla \cdot \underline{B})}_{=0 \text{ (as } \underline{v} \perp \underline{B})} - \underline{B} (\nabla \cdot \underline{v}) \\ &= 0 \text{ (no mag. monopoles)} \end{aligned}$$

$$\Rightarrow \frac{d\mathbf{B}}{dt} = (\mathbf{B} \cdot \nabla) \mathbf{v} + \frac{\mathbf{B}}{\rho} \cdot \frac{d\rho}{dt}$$

by substituting for $\nabla \cdot \mathbf{v}$

$$\rightarrow \text{then use the quotient rule: } \frac{d}{dx} \left(\frac{u}{v} \right) = \frac{v \frac{du}{dx} - u \frac{dv}{dx}}{v^2}$$

$$\Rightarrow \frac{d}{dt} \left(\frac{\mathbf{B}}{\rho_m} \right) = \frac{1}{\rho} \frac{d\mathbf{B}}{dt} - \frac{\mathbf{B}}{\rho^2} \frac{d\rho}{dt}$$

$\rightarrow$ and substitute above: (by dividing all terms by ρ)

$$\therefore \boxed{\frac{d}{dt} \left(\frac{\mathbf{B}}{\rho} \right) = \left(\frac{\mathbf{B}}{\rho} \cdot \nabla \right) \mathbf{v}} \leftarrow \text{Truesdell theorem}$$

$\frac{\mathbf{B}}{\rho} \leftarrow$ number of field lines per unit mass of the plasma

• when mass flow $\perp$ to B-field RHS vanishes

$\Rightarrow \frac{\mathbf{B}}{\rho}$ becomes conserved quantity

$\Rightarrow$ i.e. mass motion can only occur together with the B-field

$\rightarrow$ magnetic flux is "frozen in" the plasma

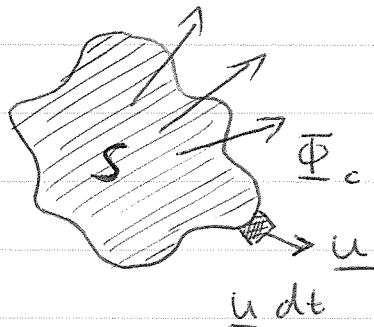
$\rightarrow$ mass flow along magnetic field lines

(in highly conductive plasmas, e.g. Solar prominences)

→ alternative derivation from Faraday's Law:

$$\underline{E} = - \frac{d\Phi}{dt} = V \quad (\text{induced voltage})$$

$$= - \oint \underline{E} \cdot d\underline{\ell}$$



$$\Rightarrow d\underline{S} = (\underline{u} dt) \times d\underline{\ell}$$

$$\text{magnetic flux: } \Phi_B = \iint_S \underline{B} \cdot d\underline{S}$$

$$\Rightarrow \frac{d\Phi_c}{dt} = \int_S \frac{\partial \underline{B}}{\partial t} \cdot d\underline{S} + \int_S \underline{B} \cdot \frac{d\underline{S}}{dt} = \text{change in } \underline{B} + \text{change in } \underline{S}$$

$$= \int_S (-\nabla \times \underline{E}) \cdot d\underline{S} + \oint (\underline{B} \cdot \underline{u}) \times d\underline{\ell}$$

$$-\frac{\partial \underline{B}}{\partial t} = \nabla \times \underline{E}$$

Faraday's Law

$$\rightarrow \text{using Stokes theorem: } \oint_{\ell} \underline{A} \cdot d\underline{\ell} = \iint_S \nabla \times \underline{A} \cdot d\underline{S}$$

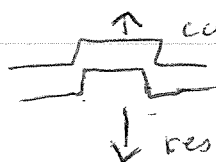
$$= - \oint_c \underline{E} \cdot d\underline{\ell} - \oint (\underline{u} \times \underline{B}) \cdot d\underline{\ell}$$

$$\Rightarrow \boxed{\frac{d\Phi}{dt} = - \oint (\underline{E} + \underline{u} \times \underline{B}) \cdot d\underline{\ell}}$$

$$\text{from Ohm's Law: } \underline{E} + \underline{u} \times \underline{B} = \eta \underline{j}$$

$$\Rightarrow \text{if } \eta \rightarrow 0 \Rightarrow \frac{d\Phi}{dt} = 0$$

i.e. flux through surface dS moving with plasma is conserved if $\eta \rightarrow 0$



→ R_m !

Magnetic pressure

vector identities:

$$\nabla(\underline{B} \cdot \underline{B}) = 2 [\underline{B} \times (\nabla \times \underline{B}) + (\underline{B} \cdot \nabla) \underline{B}]$$

$$- (\nabla \times \underline{B}) \times \underline{B} = \frac{1}{2} \nabla(\underline{B} \cdot \underline{B}) - (\underline{B} \cdot \nabla) \underline{B} = \underline{B} \times (\nabla \times \underline{B})$$

→ start from Ampère's Law: $\nabla \times \underline{B} = \mu_0 \underline{j}$

→ take curl: $\underline{j} \times \underline{B} = \frac{1}{\mu_0} (\nabla \times \underline{B}) \times \underline{B}$

$$= -\frac{1}{\mu_0} \underline{B} \times (\nabla \times \underline{B})$$

then employ vector identity: $\underline{A} \times (\nabla \times \underline{B}) = (\nabla \underline{B}) \underline{A}_c - (\underline{A} \cdot \nabla) \underline{B}$
constant in differentiation by ∇

$$\Rightarrow \underline{j} \times \underline{B} = -\frac{1}{\mu_0} [(\nabla \cdot \underline{B}) \underline{B} - (\underline{B} \cdot \nabla) \underline{B}]$$

$$= -\frac{1}{2\mu_0} \nabla(B^2) + \frac{1}{\mu_0} (\underline{B} \cdot \nabla) \underline{B}$$

analogy to convective derivative

$$\text{since } (\nabla \underline{B}) \cdot \underline{B} = \frac{1}{2} \nabla(\underline{B} \cdot \underline{B})$$

→ substitute into momentum equation: $\rho \frac{d\underline{u}}{dt} = \underline{j} \times \underline{B} - \nabla P$

$$\therefore \rho \frac{d\underline{u}}{dt} = -\nabla P - \frac{\nabla B^2}{2\mu_0} + \frac{1}{\mu_0} (\underline{B} \cdot \nabla) \underline{B}$$

thermal pressure
gradient

magnetic pressure gradient

magnetic tension force

$$\Rightarrow \text{magnetic pressure} = \frac{B^2}{2\mu_0}$$

(magnetic energy density)

$\rightarrow$ used to compress Z-pinch

$$\text{implosion stops when } \rho u^2 = \frac{B^2}{2\mu_0}$$

$$\Rightarrow \text{magnetic tension force} \sim \frac{1}{\mu_0} (\underline{B} \cdot \nabla) \underline{B}$$

in direction of curvature of the field

$\Rightarrow$ opposes bending of mag. field lines

through tension, i.e. restoring force $\Rightarrow$ wave motion

(magnetic Alfvén waves - next lecture)



$\rightarrow$ ignore if $\eta \rightarrow 0$